

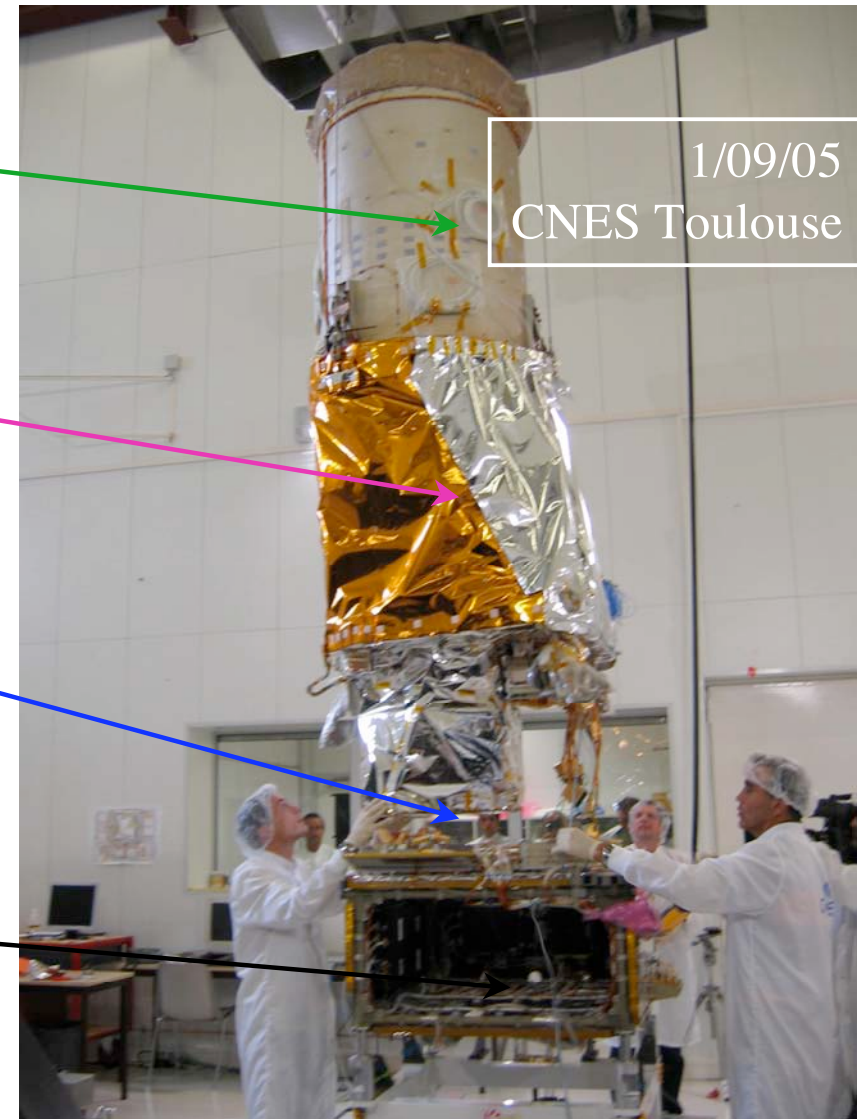
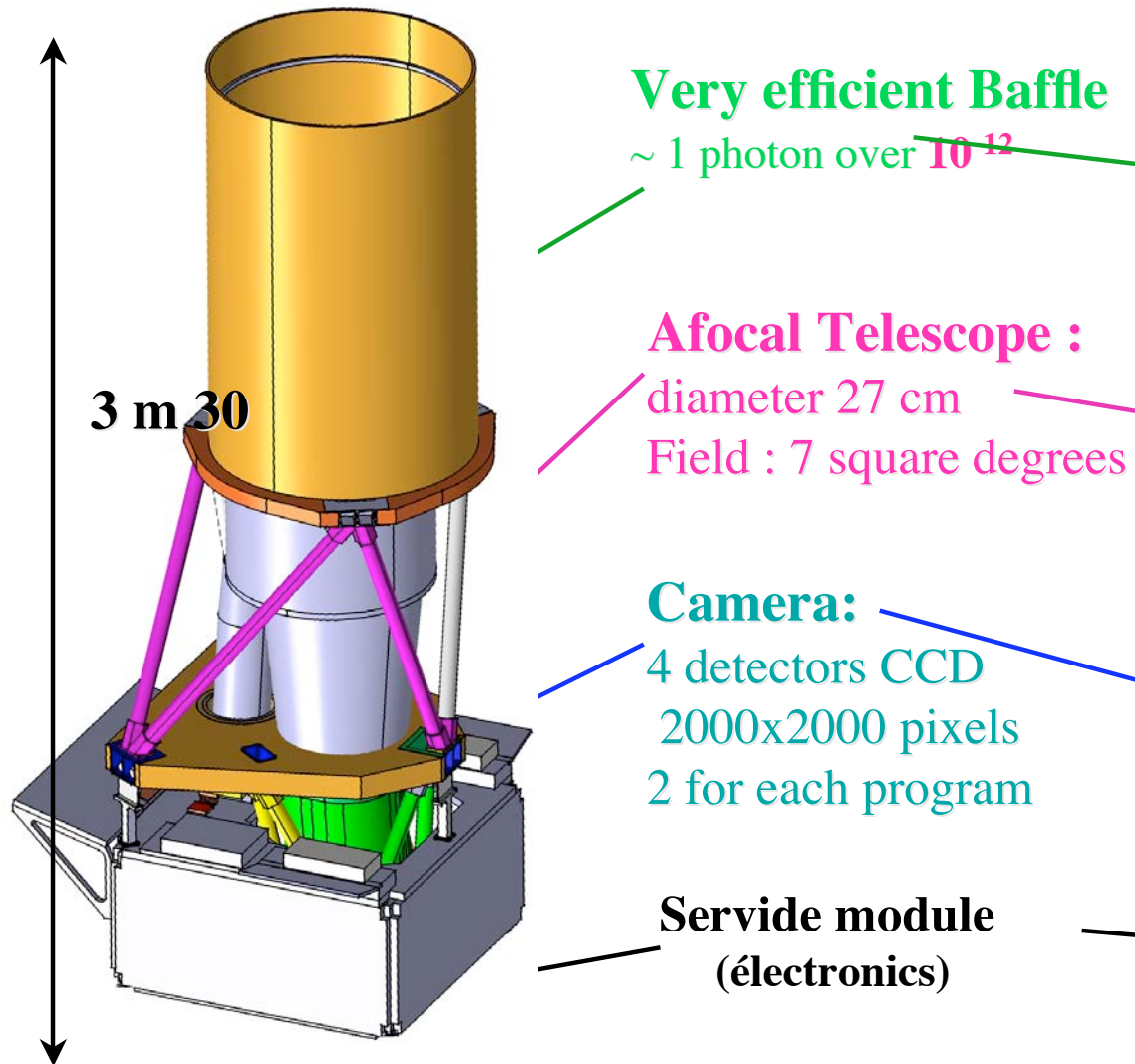
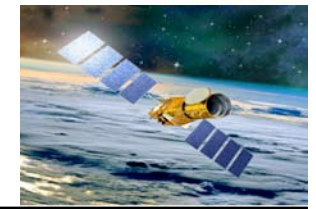
CoRoT *From stars to habitable planets*

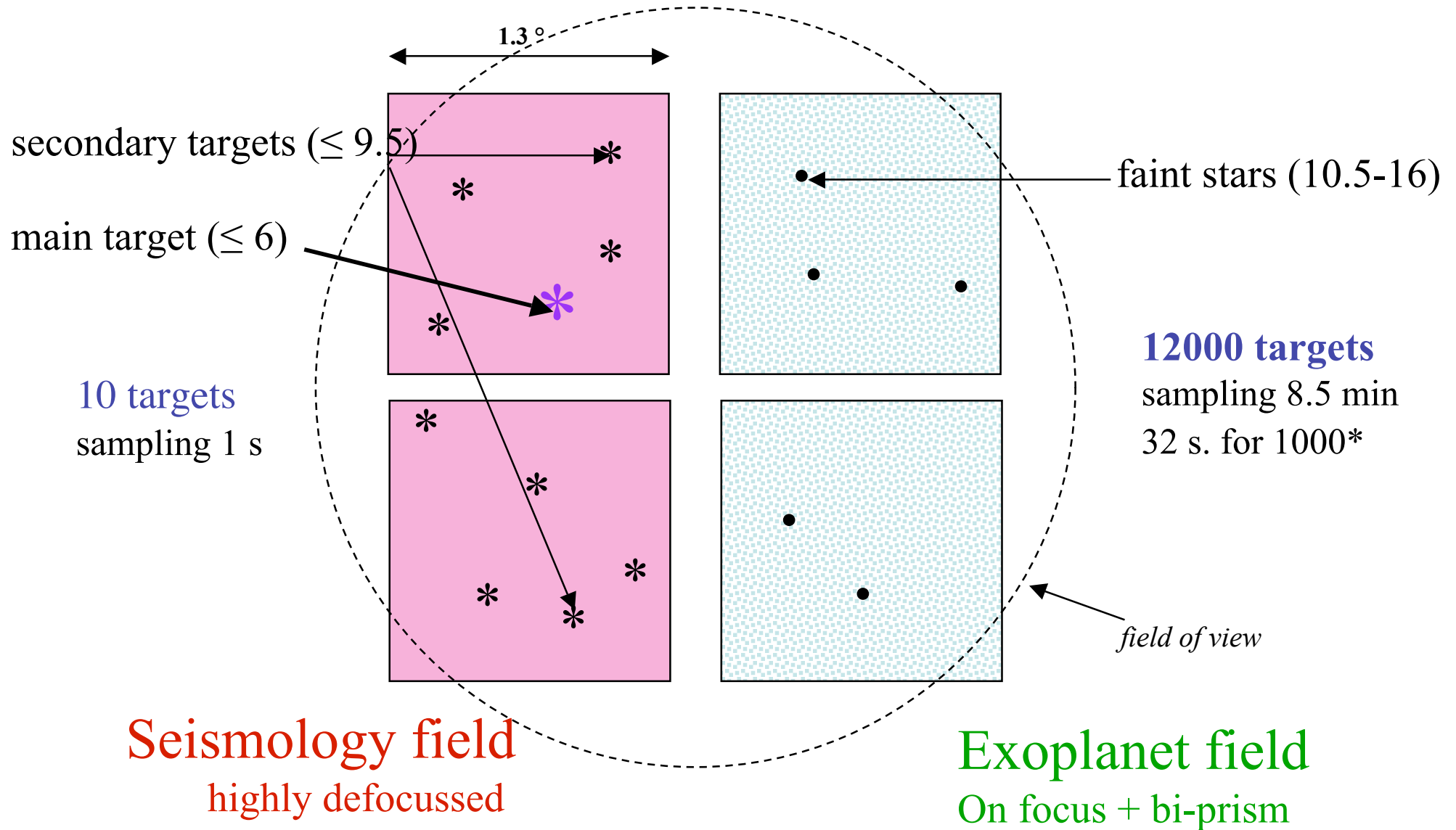
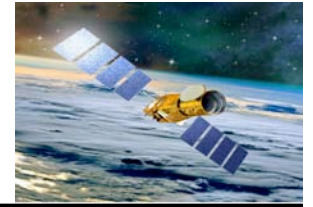
General presentation
Scientific and technical
Organisation

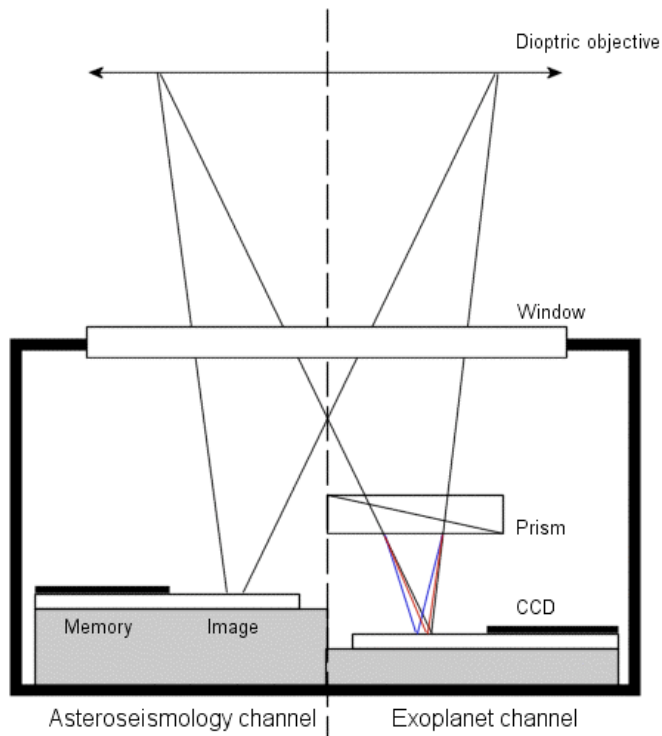
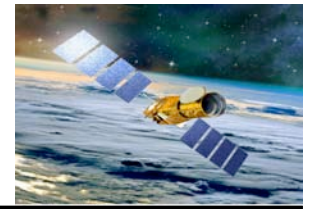
<http://corot.oamp.fr>

KASC Kick-off meeting, IAS Orsay october 29th 2007

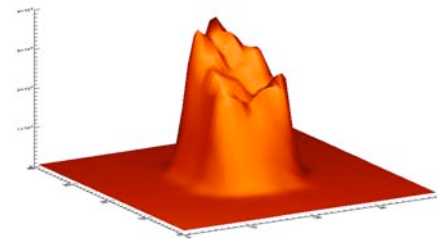
The instrument





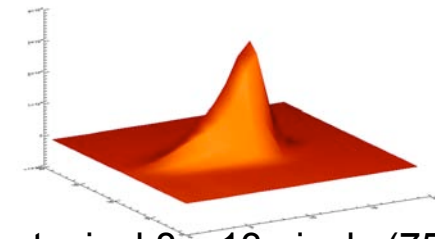


seismology



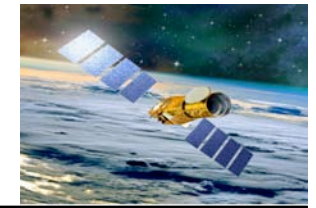
400 to 300 pixels

exoplanet



typical 8 x 13 pixels (75 pixels)

Depending on the magnitude the surface is 120 to 40 px.



$$F(t) = \Sigma [(\text{star} - \text{background}) \times \text{aperture}]$$

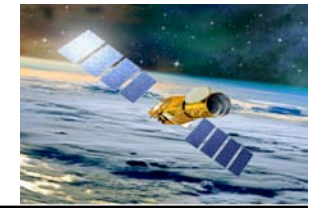
star
background
aperture

- Compute the average offset
- Compute the mean background and subtract the offset
- Substraction of the offset and background to the star image and summation of all pixels in the aperture.

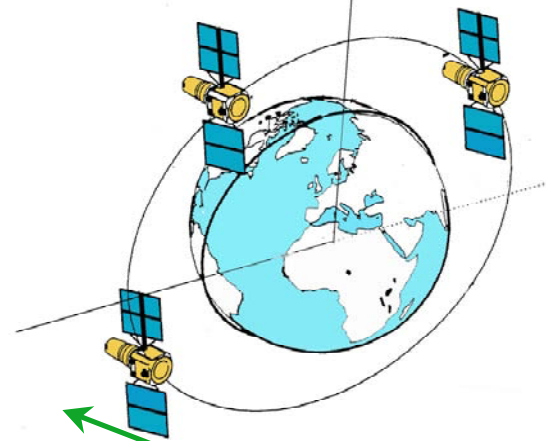
Output of the on board software: Sismofield sampling 1 s.

Exoplanet field, sampling 32 s. or 512 s. stars and background.

Offset sampling 32 s.

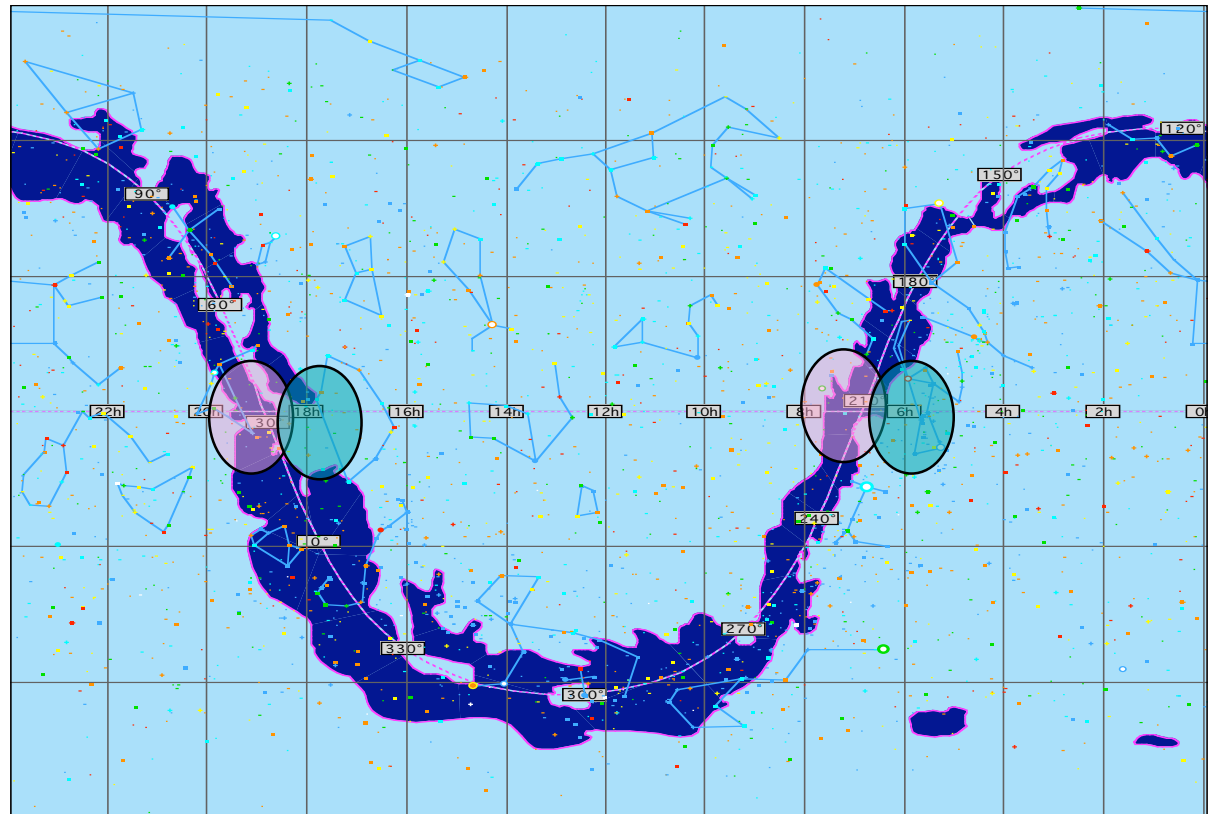


**Fixed Directions
for the whole mission**

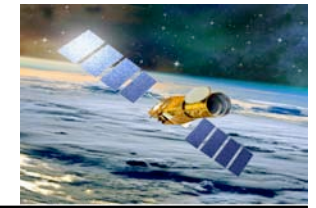


Polar Orbit

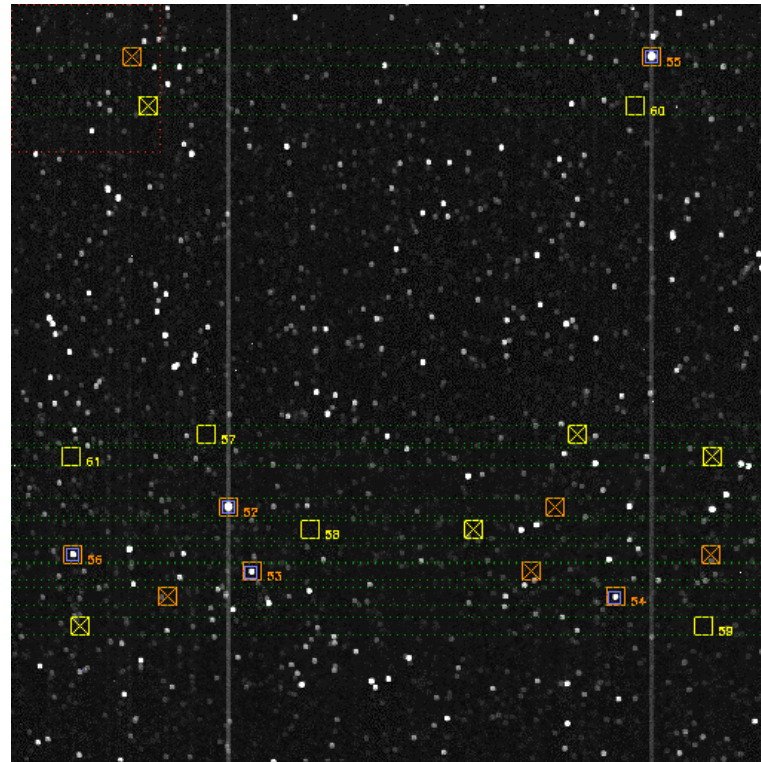
*Altitude: 896 Km,
excentricity $i=90.02$,
 $\Omega=15^\circ$,
Orbital period= 6184s.*



And drift....next year

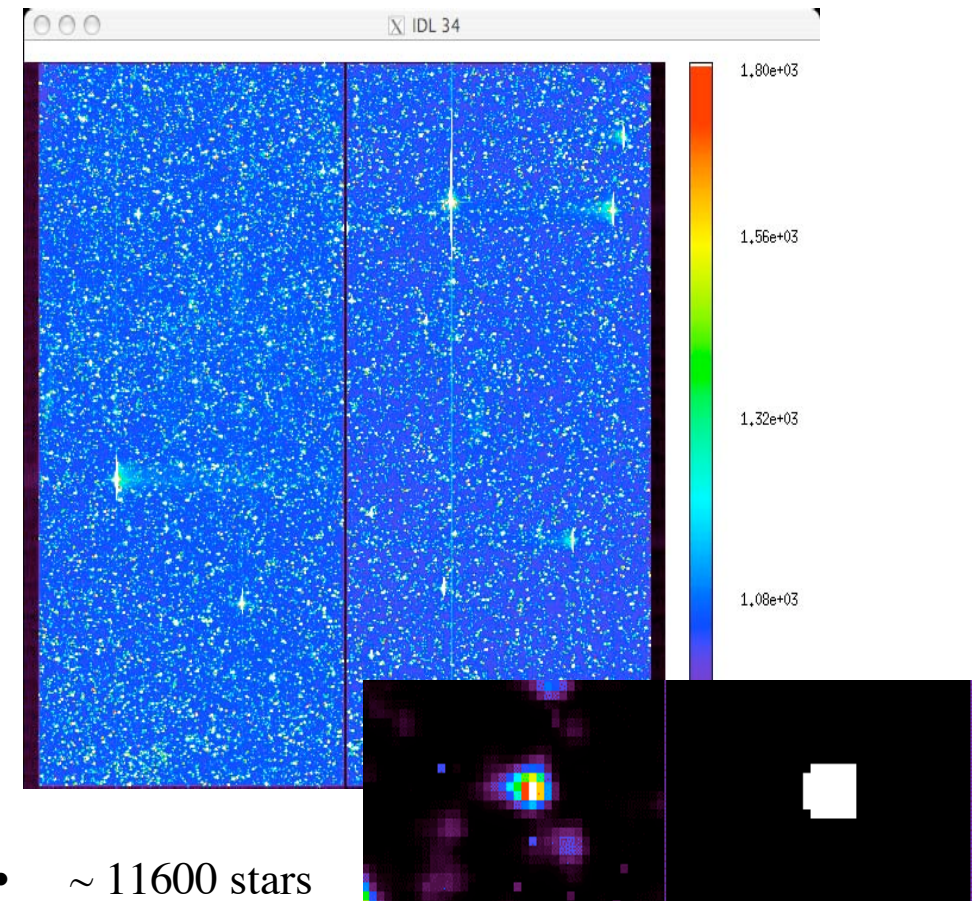


Seismology channel

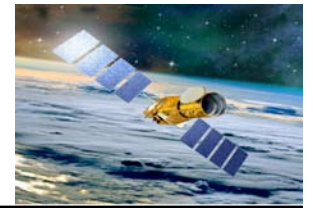


- 10 bright stars
- 10 sky background windows.

Exoplanet channel



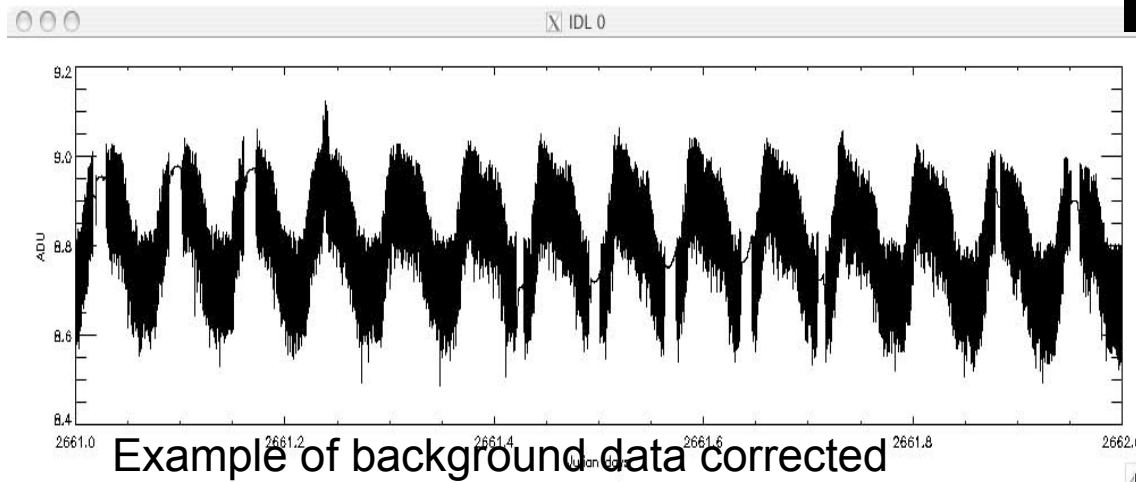
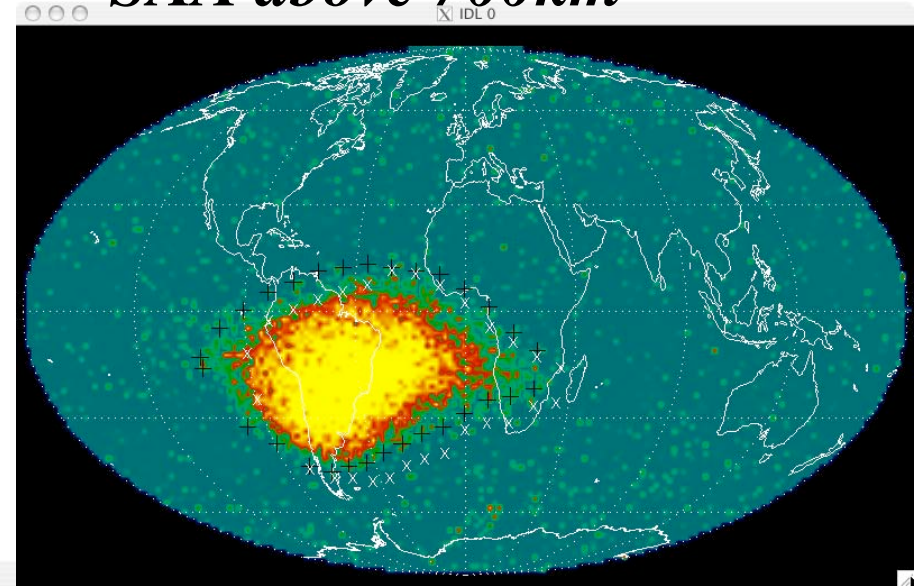
- ~ 11600 stars
- 400 sky background windows.



Runs

- IR01 80 d.
- SRc01 21d.
- LRc01 150 d.
- LRa01
- SRa01 NGC2264
-

SAA above 700km

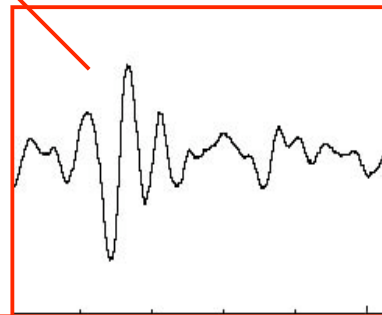
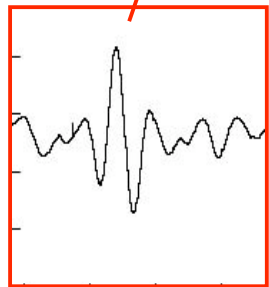
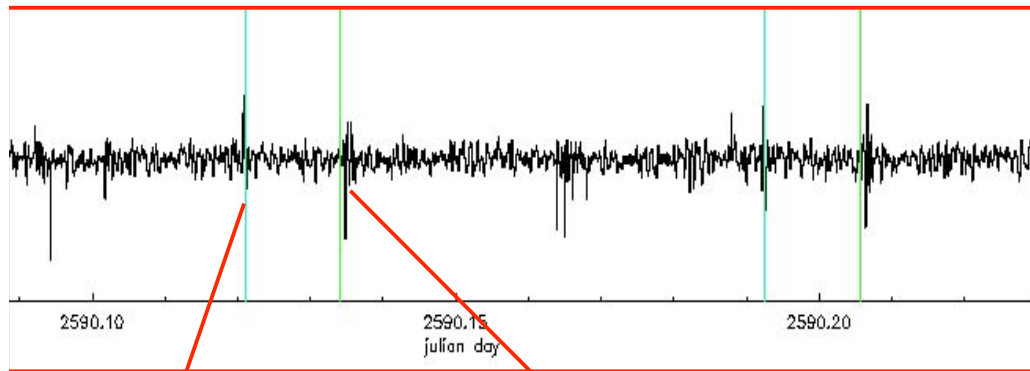
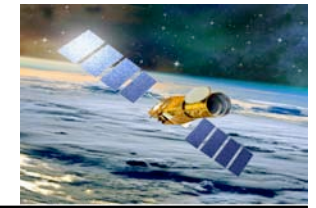


Example of background data corrected from the protons perturbations outside SAA

Data losses are mainly due to SAA crossing: 6%.

Others random events: 1 to 2%.

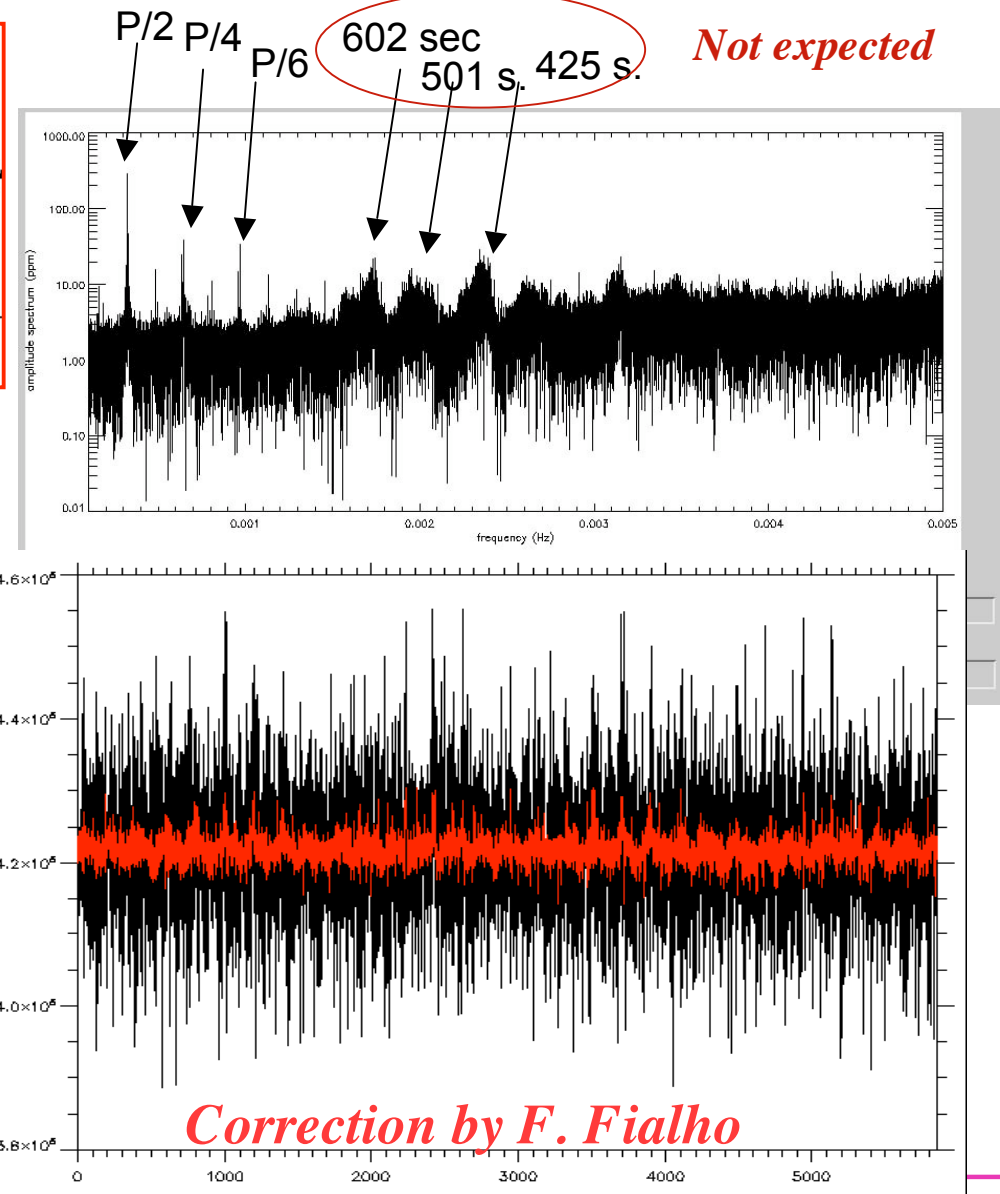
Duty cycle 92% (90 asked for).

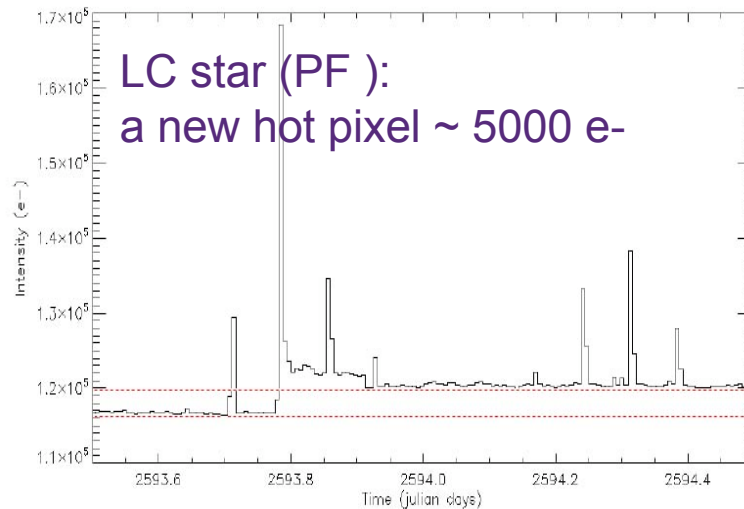
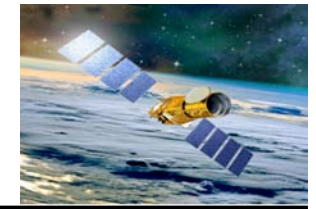


Barycentre X of a stellar image
Vibrations due to ingress/egress from eclipses

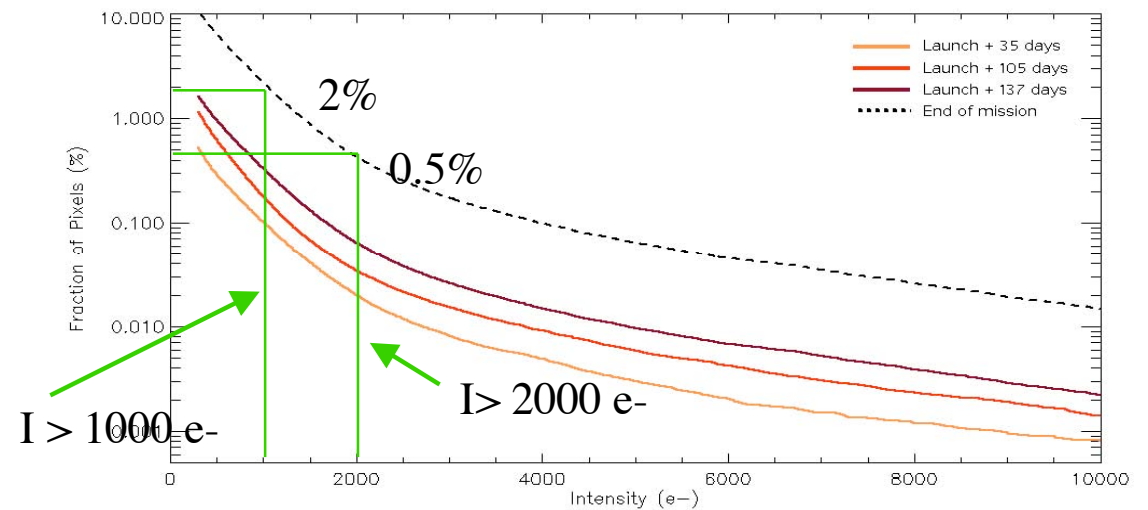
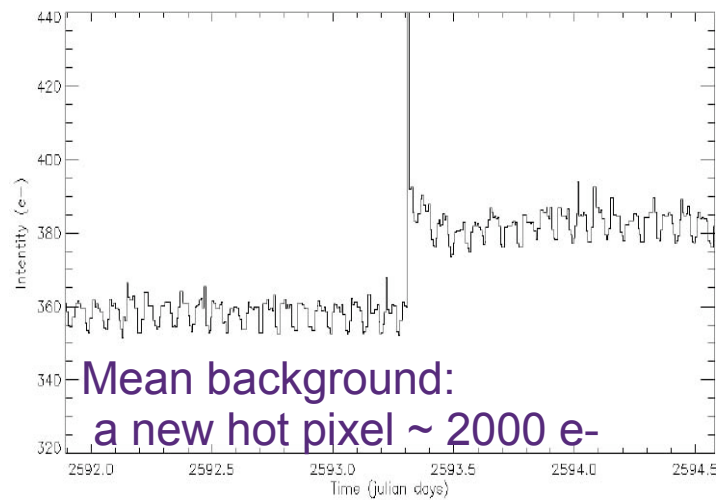
Stability Performances :

in x rms 0.12 pixel
in y rms 0.15 pixel
~ **0.3 sec. d'arc**

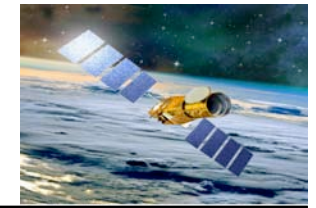




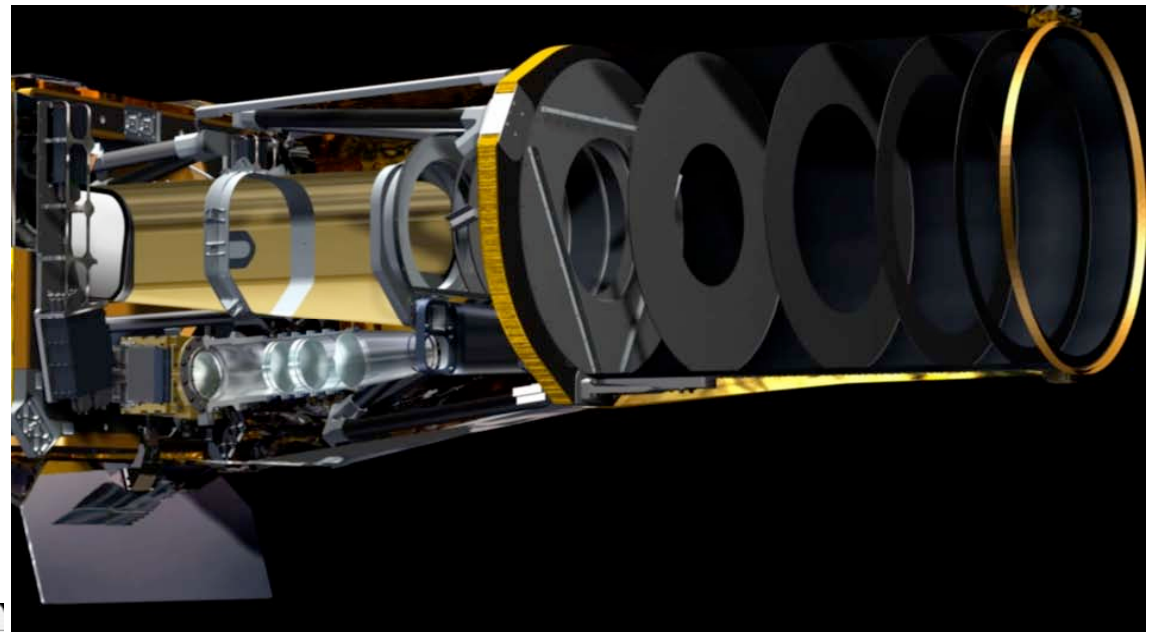
10 times more numerous than expected



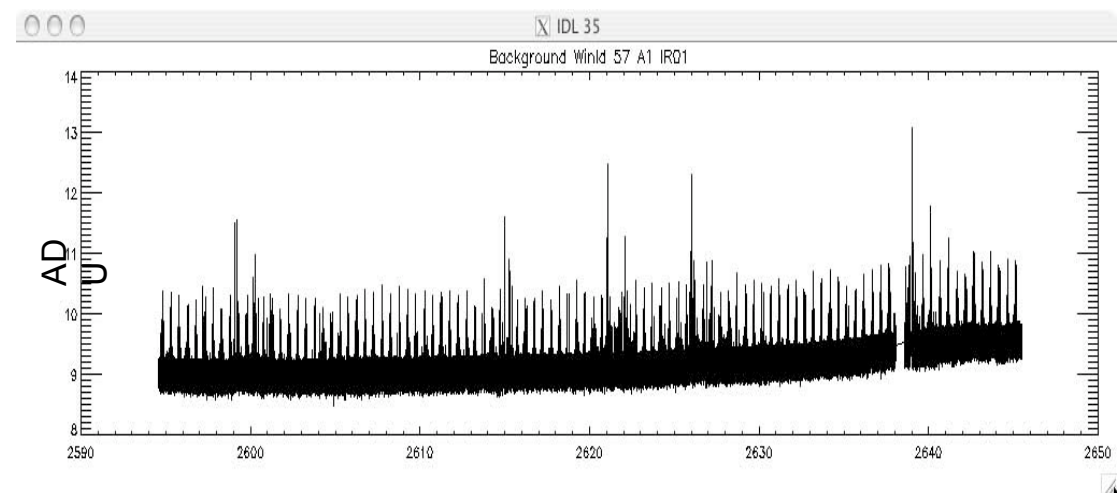
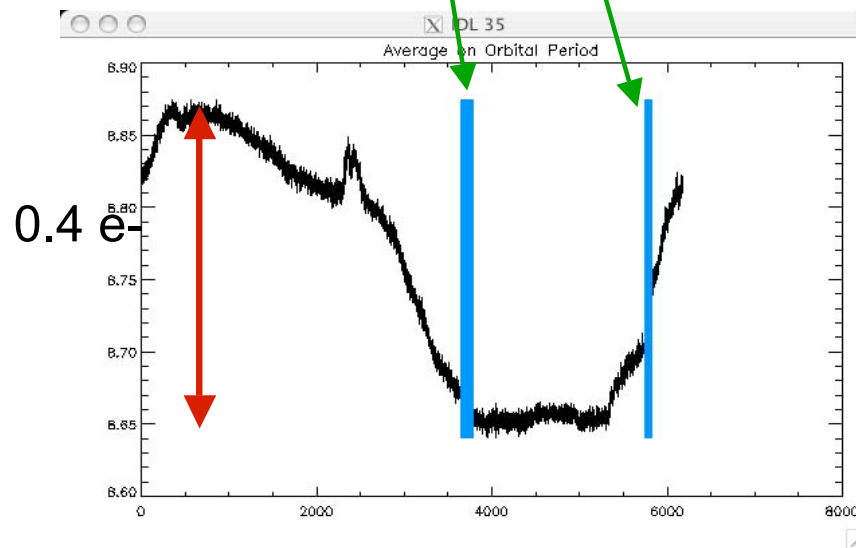
The fight against stray light

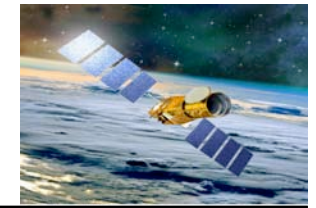


- * The afocal telescope is off-axis.
- easier to protect the focal plane against straylight
- * The baffle has two stages...



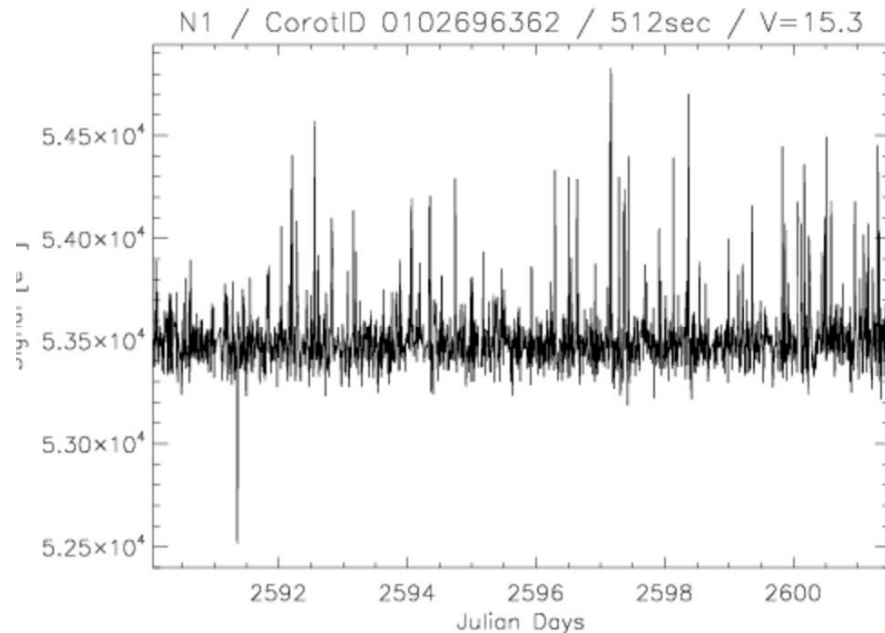
Eclipses
Ingress Egress





EXO

Spec: $7 \cdot 10^{-4}$ en 1h $m_v=15.5$



$m_v=15.4$ Integration time 1h.

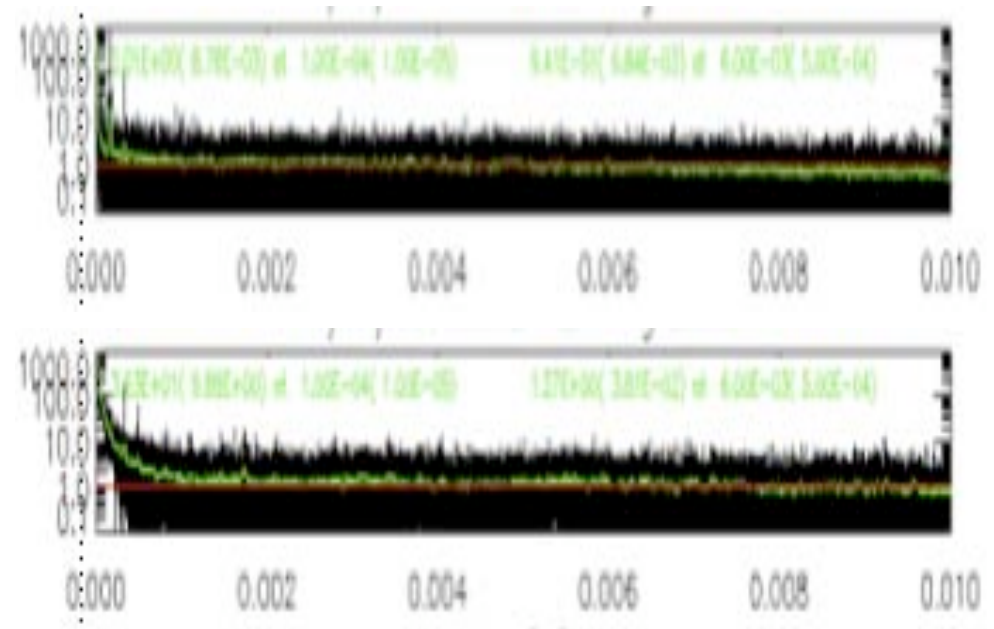
rms noise **700 ppm**

$m_v=12.3$ Integration time 1h.

rms noise **170 ppm**

SISMO

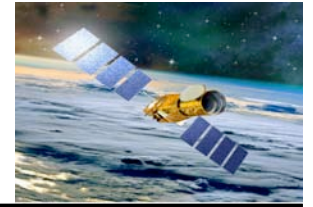
Spec: Bruit haute fréquence ($> 0.1\text{ mhz}$)
sur 5 jours, pour $m_v=5.7$ **0.6 ppm**



Presently, with very rough corrections

$m_v=5.4$, **0.8 ppm**

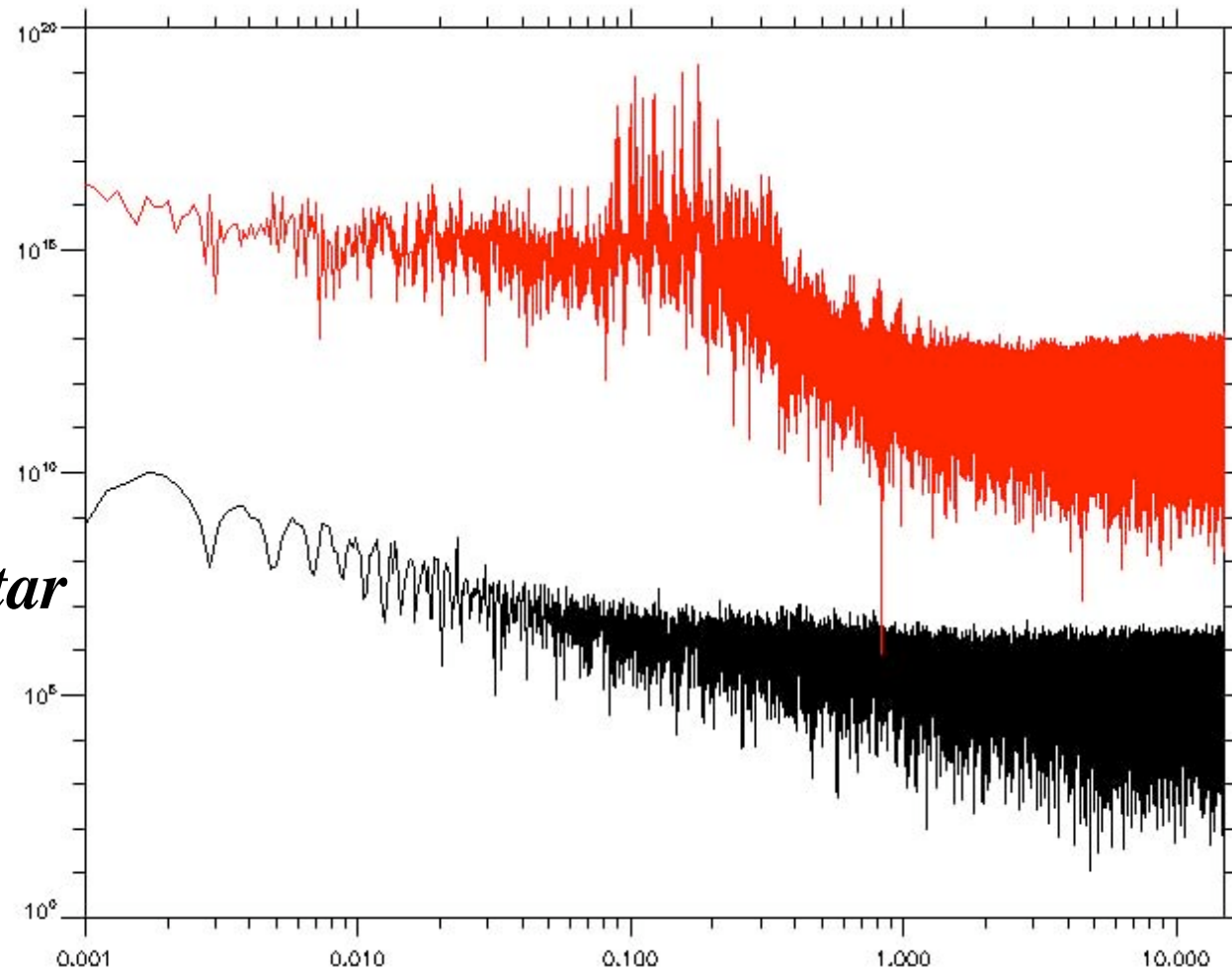
$m_v=5.7$, **0.9 ppm**

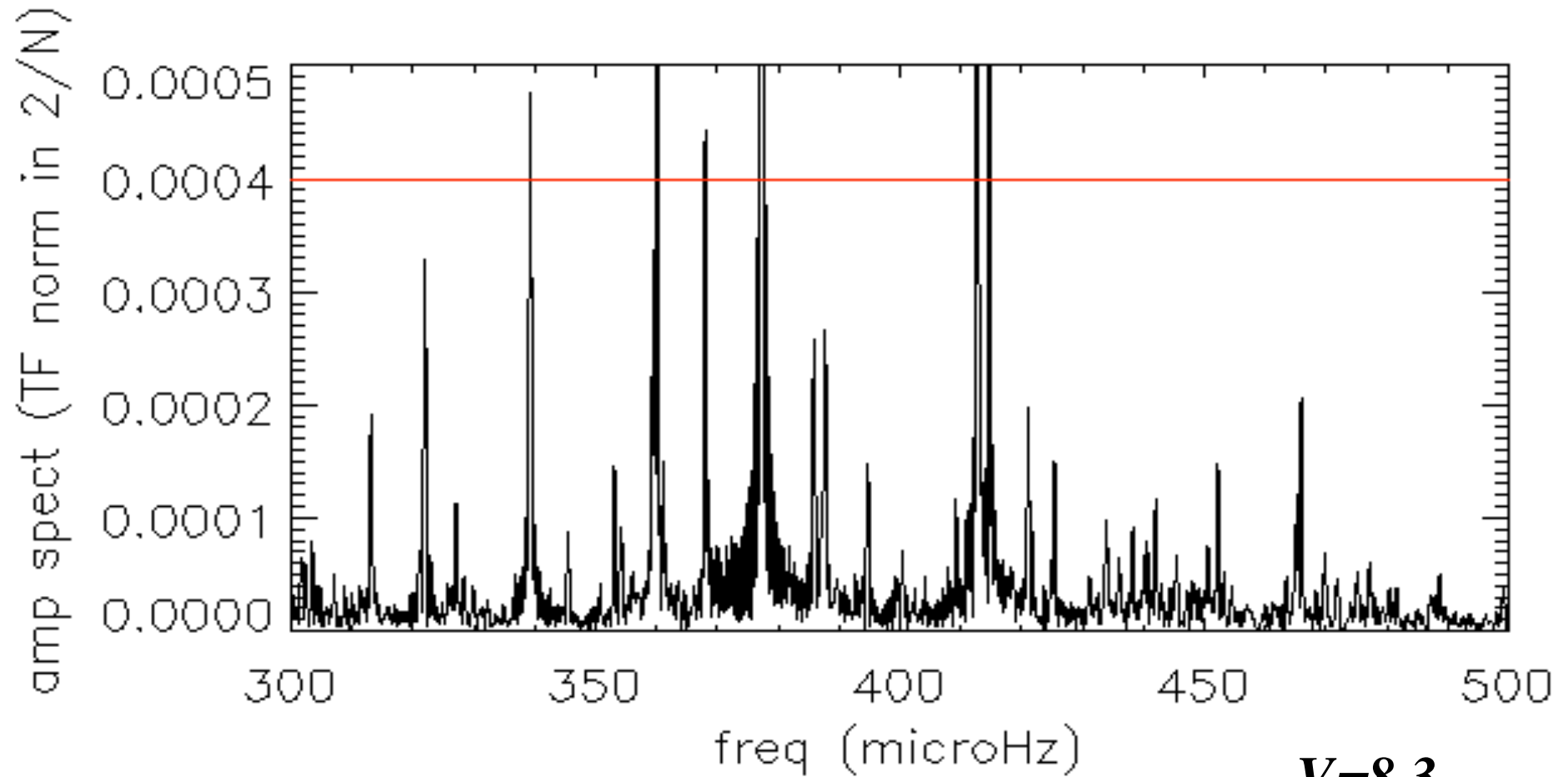
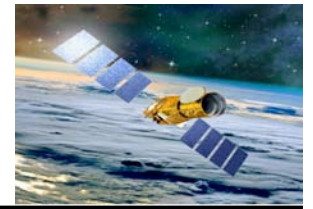


80 days

A delta scuti

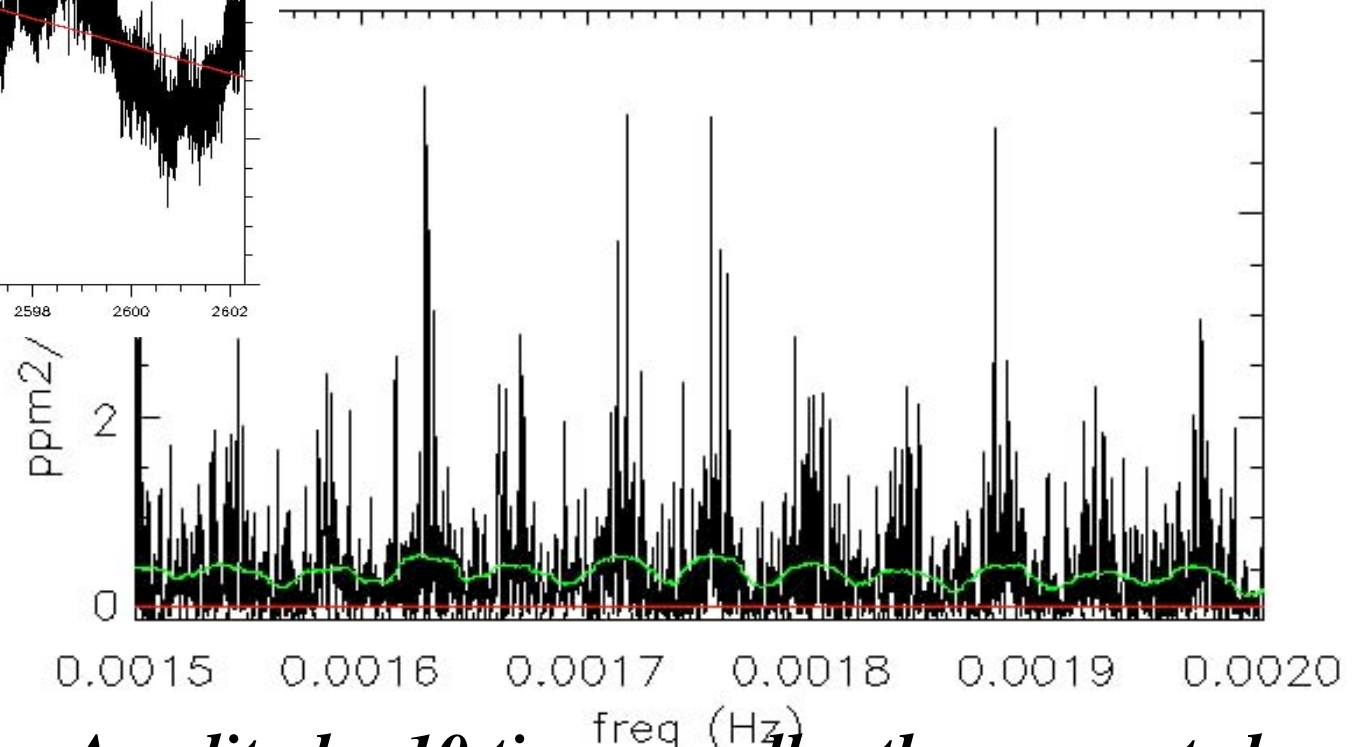
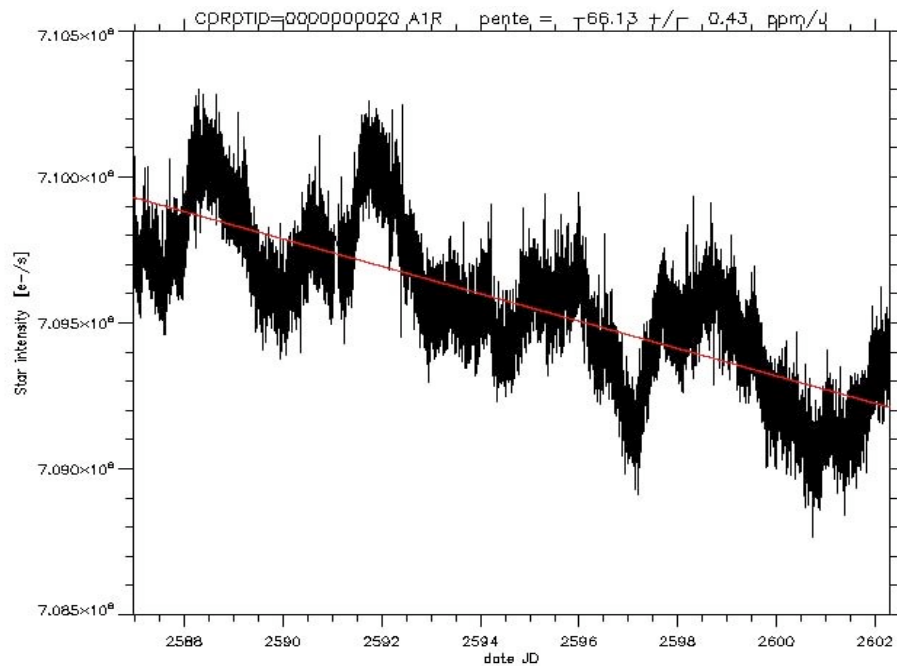
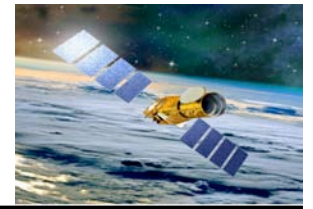
A constant (?) star



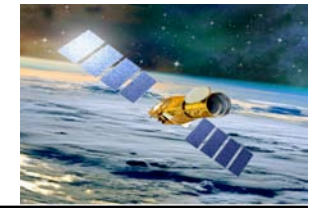


In red, present limit from the ground

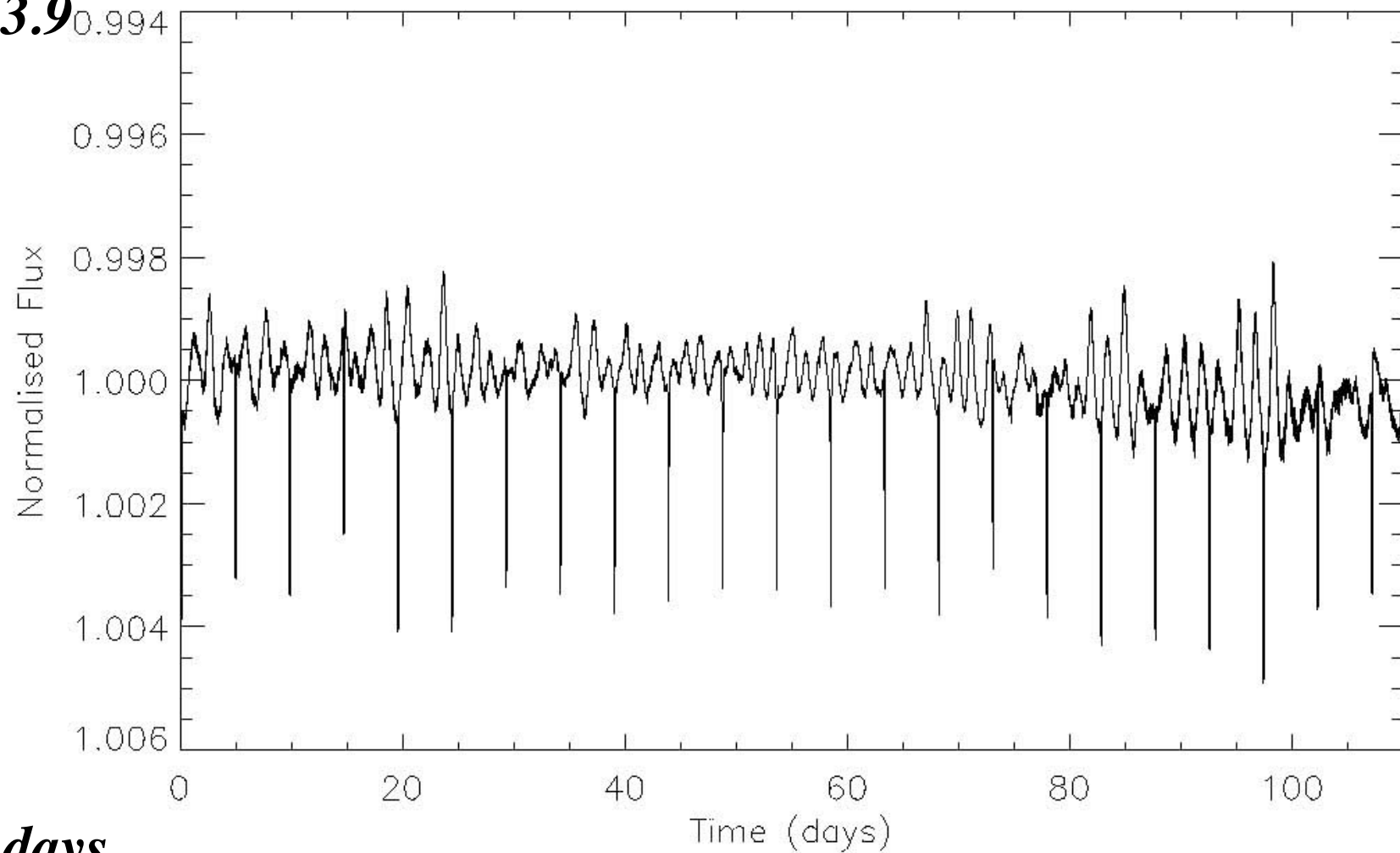
A solar like oscillator



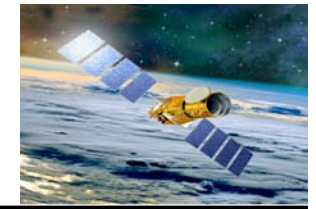
Amplitudes 10 times smaller than expected...



$V = 13.9$



120 days



Described in the CoRoT Book ESA SP 1306



**LAT, CESR / OMP
OCA
LUTH, GEPI / OPM**



Austria



Brazil



Spain



Germany

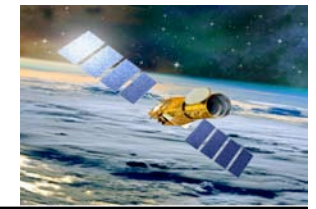


Belgium



ESA, Sc prog.

<http://corot.oamp.fr> <http://smc.cnes.fr>



Scientific Committee

In charge of

- the scientific specifications
- the programme of observations
- the data distribution policy
- the list of Co-Is and GIs

CORE programme

Seismology in seismology field

Exoplanet search in exoplanet field

Seismology
W.G.
E. Michel

Seismology
G.B. obs.
C. Catala

Exoplanet
W.G.
P. Barge

Exoplanet
GB obs.
M. Deleuil

Science Teams
Co_Is

Restricted
to CO-Is

Access to all data

Additional Programmes

Any science case outside the Core programme

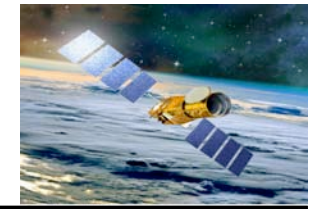
W. Weiss

Open programme
Through AOs

Sc. Teams
GIs



Co-Is, GIs and Associated Scientists



*Co-Is 5 per contributing country
majority of frenchs*

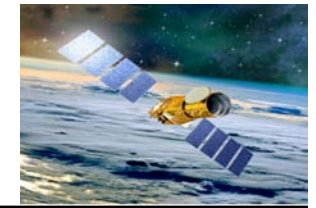
72

GI s successful proposers to Aos of Aps.....not limited

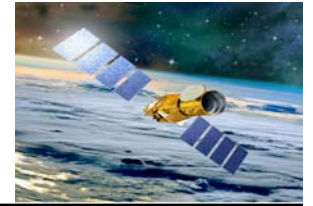
*Associated scientists: participating to a team lead by a Co-I or a GI
by the SC*



The four levels of CoRoT data



N0	Time-stamped raw data	CNES
N1	Data corrected from main instrumental effects	Science archive then LTAs <i>Upon request</i>
N2	Scientific labelled data (SC)	Science archive then LTAs
N3	Standard scientific products (SC)	Science archive then LTAs



The Scientific policy and Data Rights Document

COROT.DESPA.01.014

version 6.1

September 14th 2007

Available at:

<http://corot.oamp.fr/scomite.html>

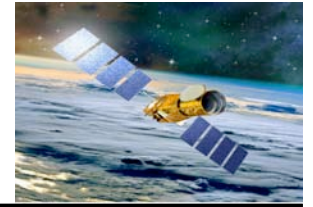
	Titre du document : Scientific policy and data rights	Référence : COROT.DESPA.01.014
	Auteurs : Annie BAGLIN	Version : 4 Date : 26/05/06 Page : 1

Scientific policy and Data rights

Préparé by par:	Annie BAGLIN	
Accepted by par:	Scientific Committee	
Pour application:		

HISTORIQUE DES MODIFICATIONS DU DOCUMENT

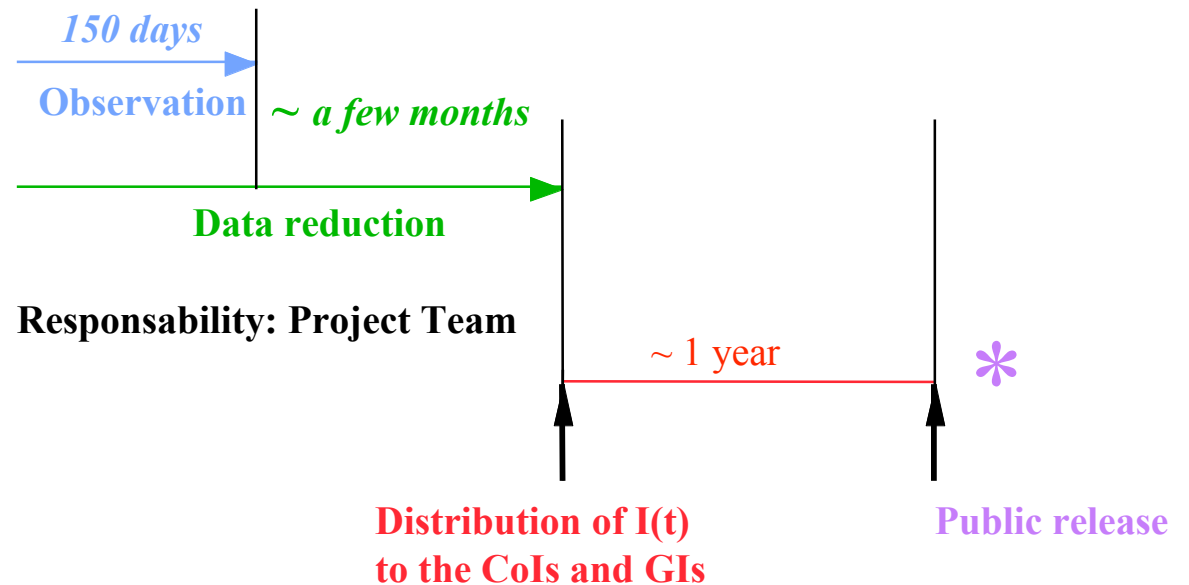
Ed.	Rév.	Date	Modifications/	Visa
1		12/02/01		
2		19/05/04	Data rights, project team, 5 WG	
3		15/03/05	Definition of CoRoT-Is and associated scientists Rights and duties of the CoRoT SC advisory board Publication policy	
4		20/05/06	3.1 of In case of reobservation of the same field	



*The data are processed
by the CMC (CNES/Labs) first levels
by the CDC (Labs) final level.*

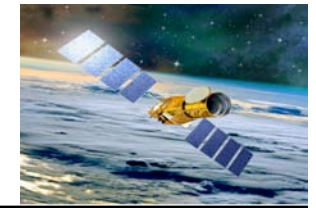
They are archived at IAS

*Long term archives
at CDS and LAEFF*





Who will get what? CoRoT programme



The Co-Is access **all** the data

The Co-Is are responsible for the **fair** use of the data
during the proprietary period

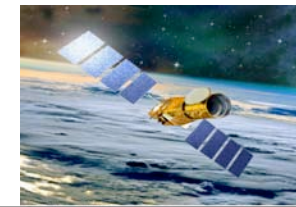
The Co-Is can analyse and interpret the data,
with the **Team** they are leading

The Team has to be known and accepted by the SC
(Co-Is and Associated Scientists)
has to comply with the publication policy

The answer to the **Call for Letter of Intent** organises this!



Who will get what? Additionnal Programmes



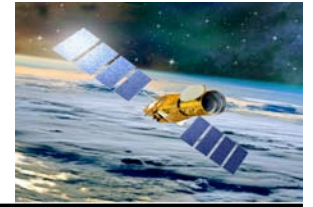
Programmes submitted through AOs
accepted by the SC
agreement on the Team

The team leader becomes GI
access the data as decided by the SC
is responsible for the fair use of these data
during the proprietary period

The members of the team (if not already Co-I or GI)
become Associated scientists for this programme



Ground based Data



COROTSKY Data Basis (OMP Toulouse):

see S. Charpinet at CW9 <http://corot.oamp.fr> Users interface at OMP

- * SISMOSKY : $V < 9.5$ in the eyes of CoRoT (13000*)
≠ parameters + Adopted
- * EXOSKY : $R < 16$ in selected area of the eyes (from EXODAT)
Photometric observations+ parameters
- * AP-EXO : selected sources
- * USNO_A2

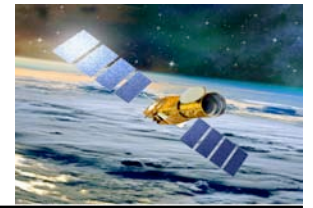
Part of the Corotsky data on observed targets will be included in N1 and N2
and can be used for requests

GAUDI (LAEFF): spectroscopic and photometric data of sismo targets (1500*)
1 year proprietary period, then public
Through the Corotsky tool or directly

EXODAT (LAM, Marseille): ground based observations + parameters
of selected fields for exoplanets (on request)

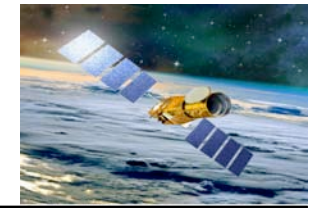


<http://www.obs-nice.fr/CoRoTpub>





Avec...la CoRoT Bossa



Myriam Adda, Conny Aerts, Georges Alecian, Claudia Andreis, Thierry Appourchaux, Armand, Bernard Authier, Michel Auvergne, Annie Baglin, Pierre Barge, Caroline Barban, Philippe Baron, Frédéric Baudin, Etienne Baudino, Thierry Beaufort, Remy Bellenger, Pernelle Bernardi, Stéphane Berrivin, Gabrielle Berthomieu, Jean-Lucien Boit, Pierre Bodin, Laurent Boissard, Cédric Bombail, Frédéric Bonneau, Pascal Bordé, Patrick Boumier, Jean-Tristan Buey, Damien Cailliau, Claude Catala, Rafael Cautain, Stéphane Charpinet, Sylviane Chaintreuil, Michel Combes, Eduardo Costa da Silva, Vincent Costes, Josiane Cuvilo, Pascal Dargent, Jean Claude David, Michel Decaudin, Celine Defay, Magali Deleuil, Alain Docclo, Rachel Drummond, Jean-Bernard Dubois, Olivier Dupuis, Gérard Esptein, Fabio Favata, Fabio Fialho, Claudie Fauré, Didier Ferrand, Bernard Foing, Michael Fridlund, Alain Gaboriaud, Rafael Garcia, Rafael Garrido, Marie-Jo Goupil, Loïc Gueguen, Cecil Gueriau, Tristan Guillot, Emmanuel Grolleau, Rachel Gruneisen, Olivier Hagogue, Misha Haywood, Anne-Marie Hubert, Gérard Huntzinger, Constance Imad, George Isaack, Eduardo Janot-Pacheco, Garry Jeanville, Farid Karioty, Laurent Kerjean, Laurent Jorda, Patrice Journoud, Martine Jouret-Perl, Vincent Lafaille, Laurent Lafond, Thien Lam-Trong, Philippe Landiech, Vincent Lapeyrere, Philippe Laudet, Nicolas Lautié, Yveline Lebreton, Hélène Le Cann, Alain Léger, Jean-Pierre Lefaure, Philippe Lemaire, Liliane Leporati, Patrick Levacher, Antoine Llebaria, Yiu-Yin Longval, Alain Magnan, André Mangeney, Philippe Mathias, Dominique Mazari, Jean-Michel Ménager, Eric Michel, Jean-Pierre Michel, Jean-Pierre Moalic, Benoît Mosser, Claire Moutou, Raymond Morat, Coralie Neiner, Arlette Noels, Marc Ollivier, Jean-Luc Orcesi, Jérôme Parisot, Véronique Pascal, René Perez, Gispert Peter, Alain Piacentino, Leonardo Pinheiro, Philippe Plasson, Jacqueline Platzer, Ennio Poretti, Dominique Pouliquen, Françoise Praderie, Janine Provost, Pascal Puget, Fy Rakotoarivelo, Heike Rauer, Jean-Louis Reynaud, Noël Robichon, Renaud Romagnan, Christine Rossin, André Roussin, Daniel Rouan, Ian Roxburgh, Michel Saisse, Reza Samadi, Régis Schmidt, Jean Schneider, Noah Schwartz, Isabelle Sebbag, Guy Simon, Enrique Solano, Manfred Steller, Magali Tello, Linda Tomasini, Jacqueline Thouvy, Didier Tiphène, Dominique Toubanc, Sylvaine Turck-Chièze, Gérard Vauclair, Sylvie Vauclair, Olivier Vandermarcq, Claude van't Veer, Jean-Claude Vial, Thierry Viard, André Vuillemin, Werner Weiss, Philippe Zanatta,.....et tous les autres.....
Certains qui ont disparu..... et d'autres qui sont allés vers d'autres horizons.....