

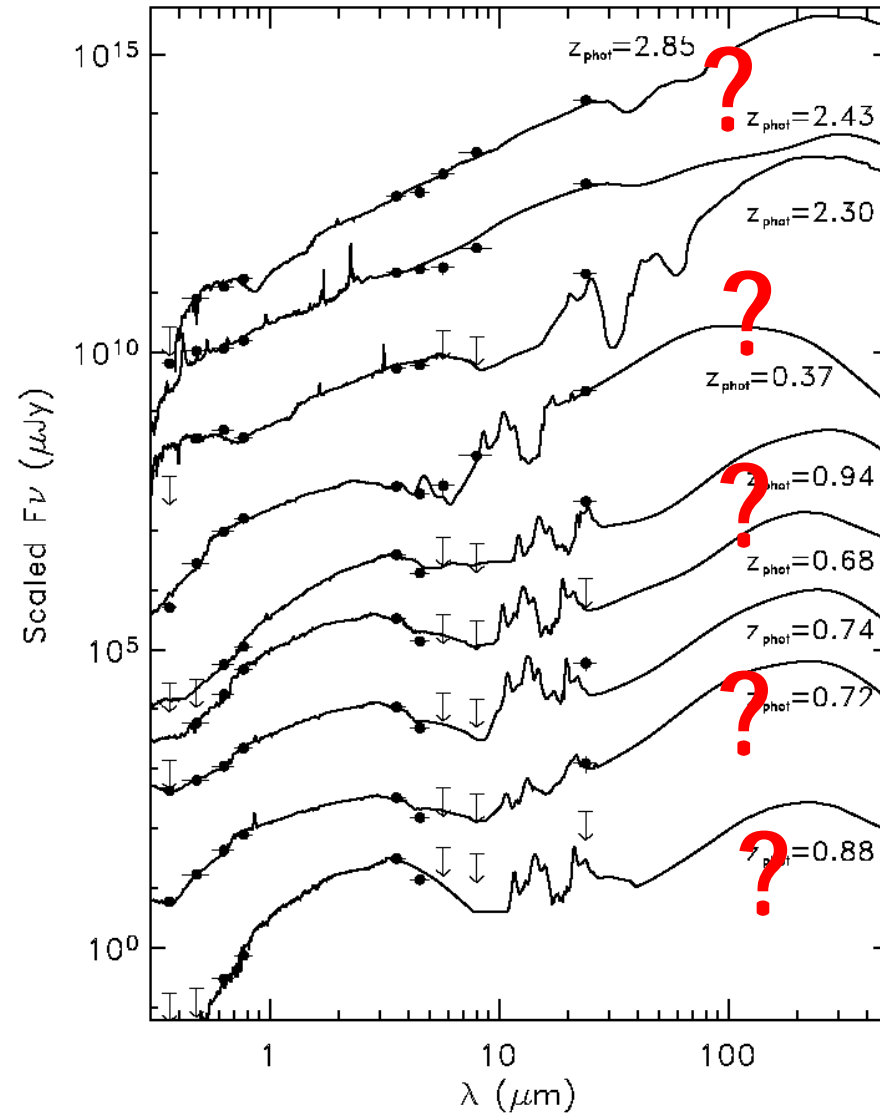
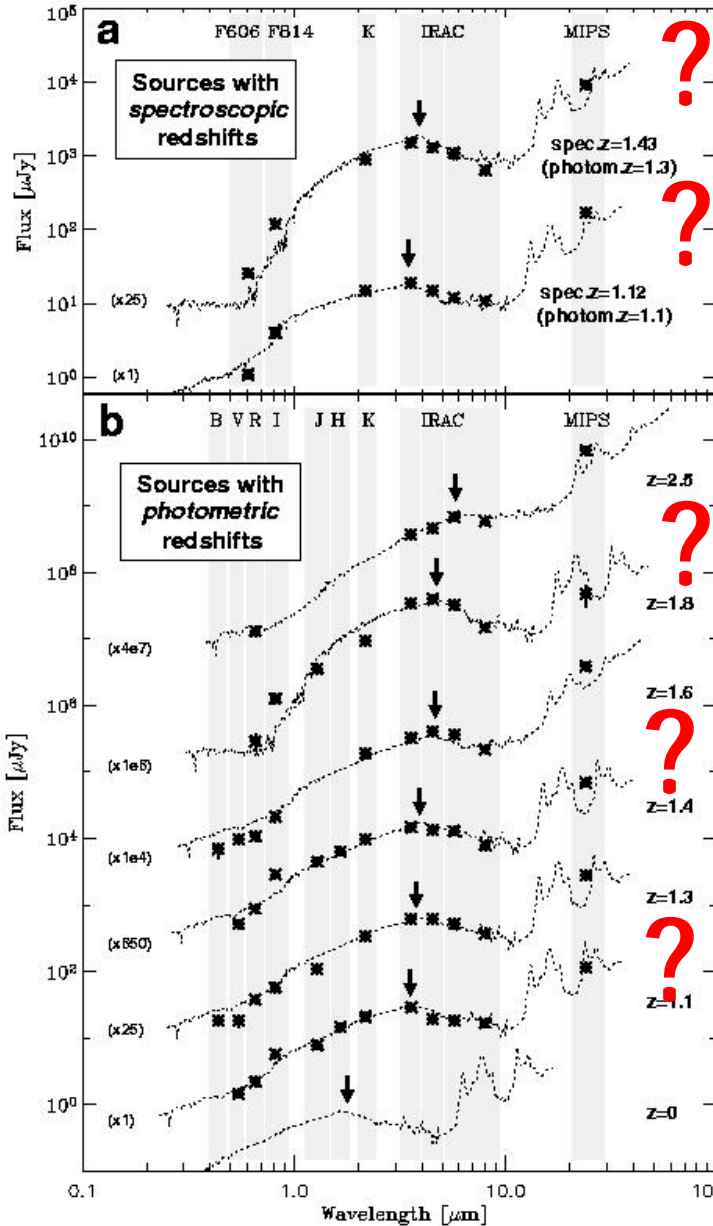


Spitzer and the Cosmic Infrared Background

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[w/ K. Caputi, G. Lagache, J-L. Puget et al.]
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Getting SEDs



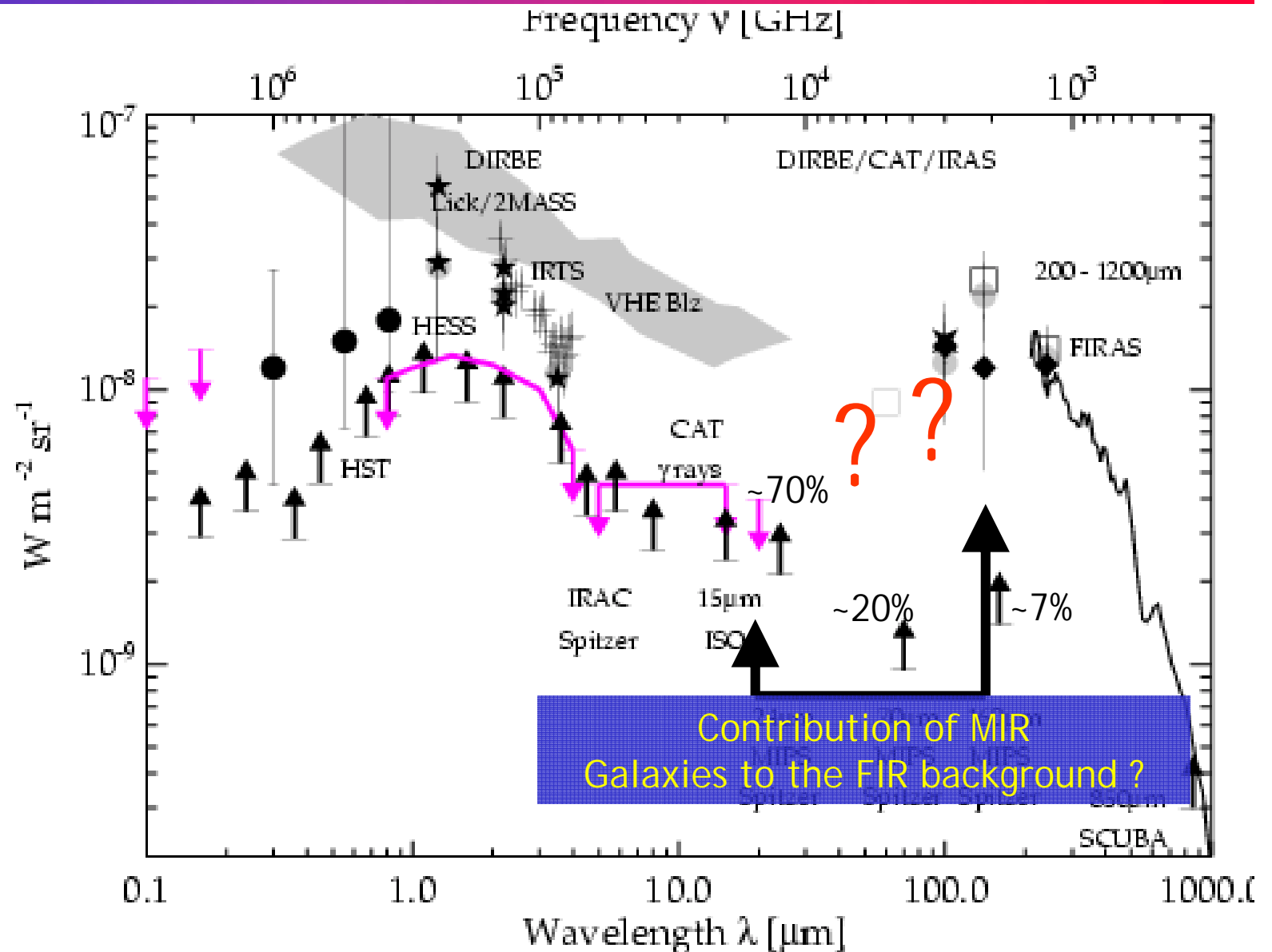
Le Floc'h et al., 2004
Lonsdale et al., 2004

Cosmic Infrared Background

-What is the Nature of the Galaxies Making-up the FIR CIB ?

-What is the Contribution of the MIR Galaxies to the FIR CIB ?

-Use 24um Galaxies to probe the CIB

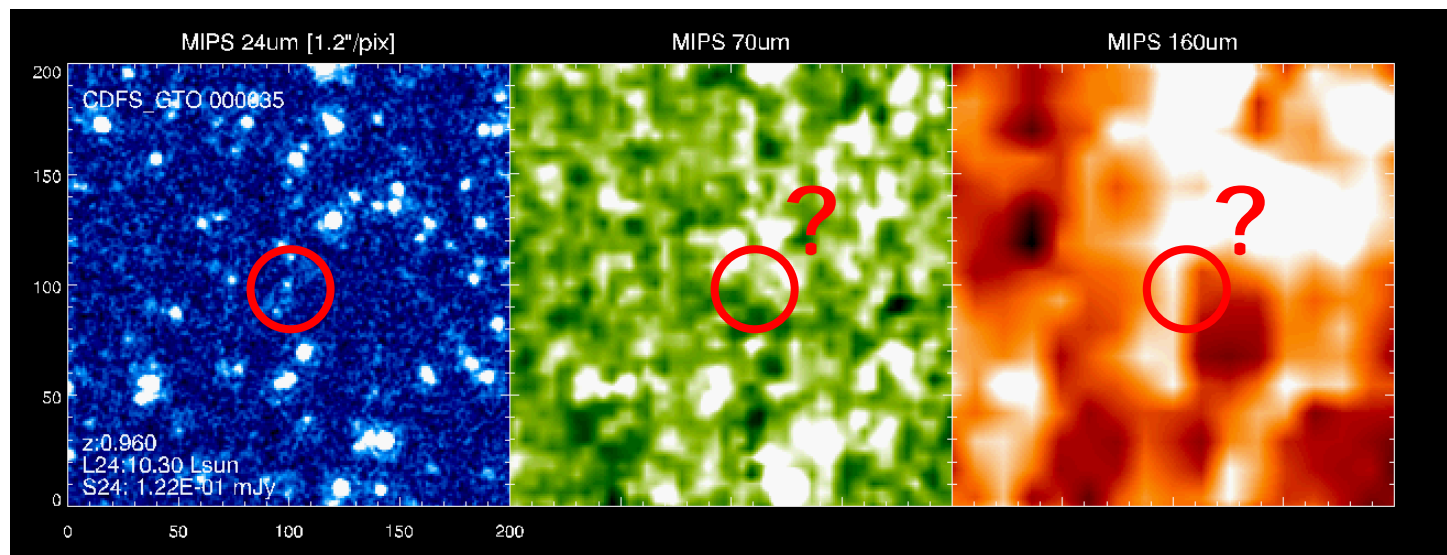
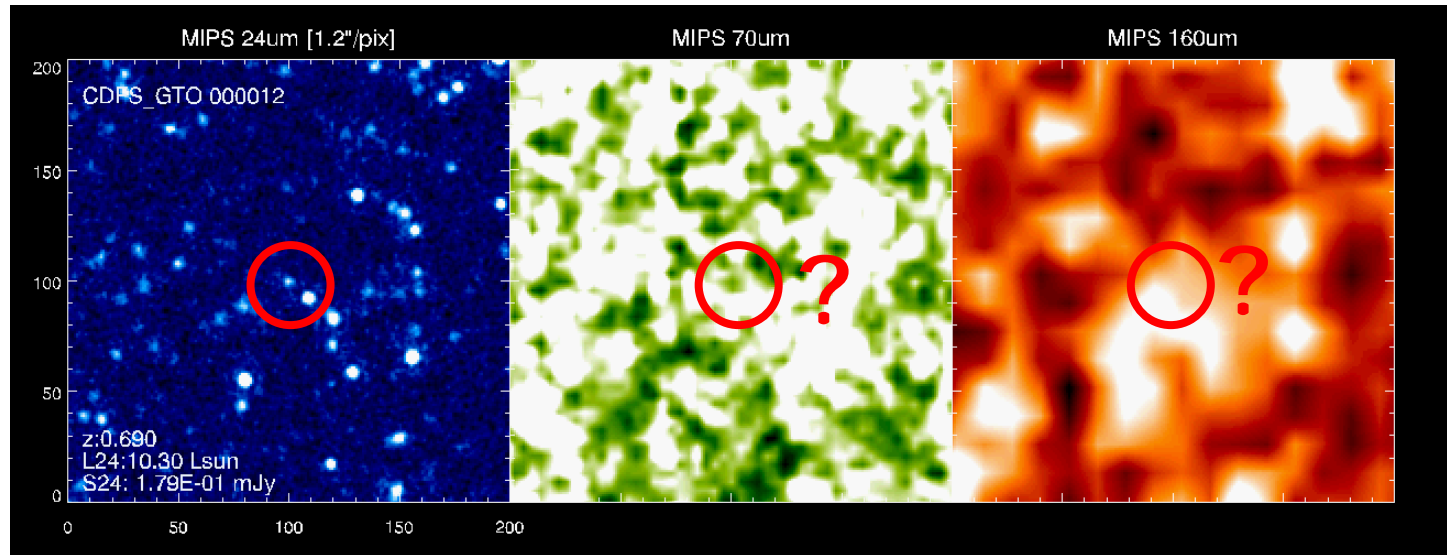


The Problem: Confusion & Noise

What to Measure
at 70 and 160
when a 24 μ m
source exists ?

Example of 24 μ m
sources: nothing
can simply be
detected in the
FIR.

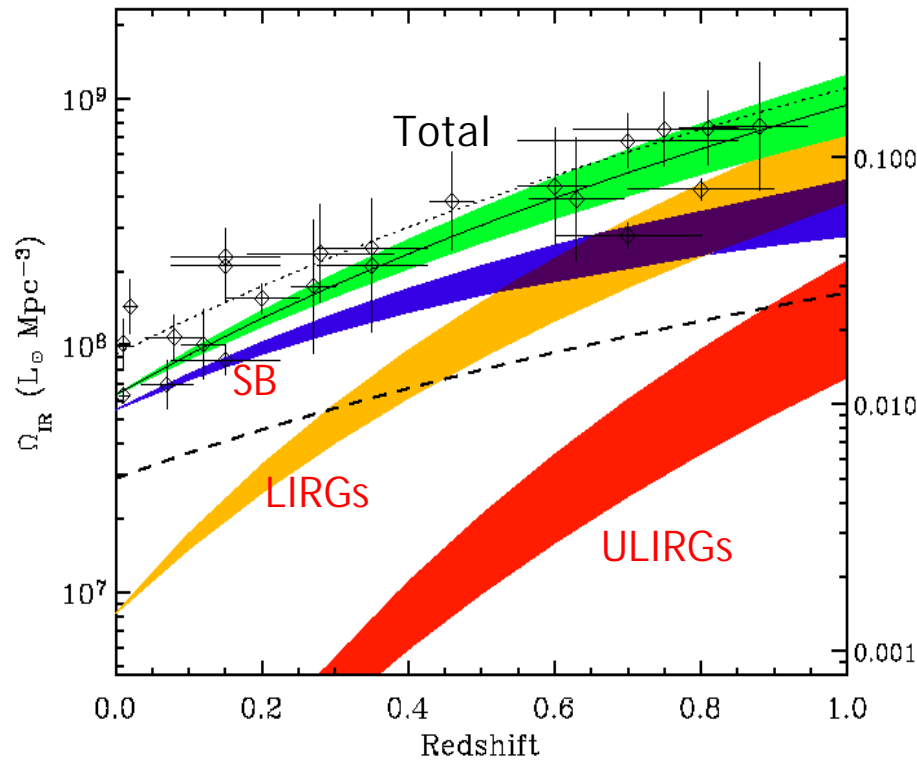
z~0.7 180 μ Jy
z~1 120 μ Jy



The Solution: 24um Sources

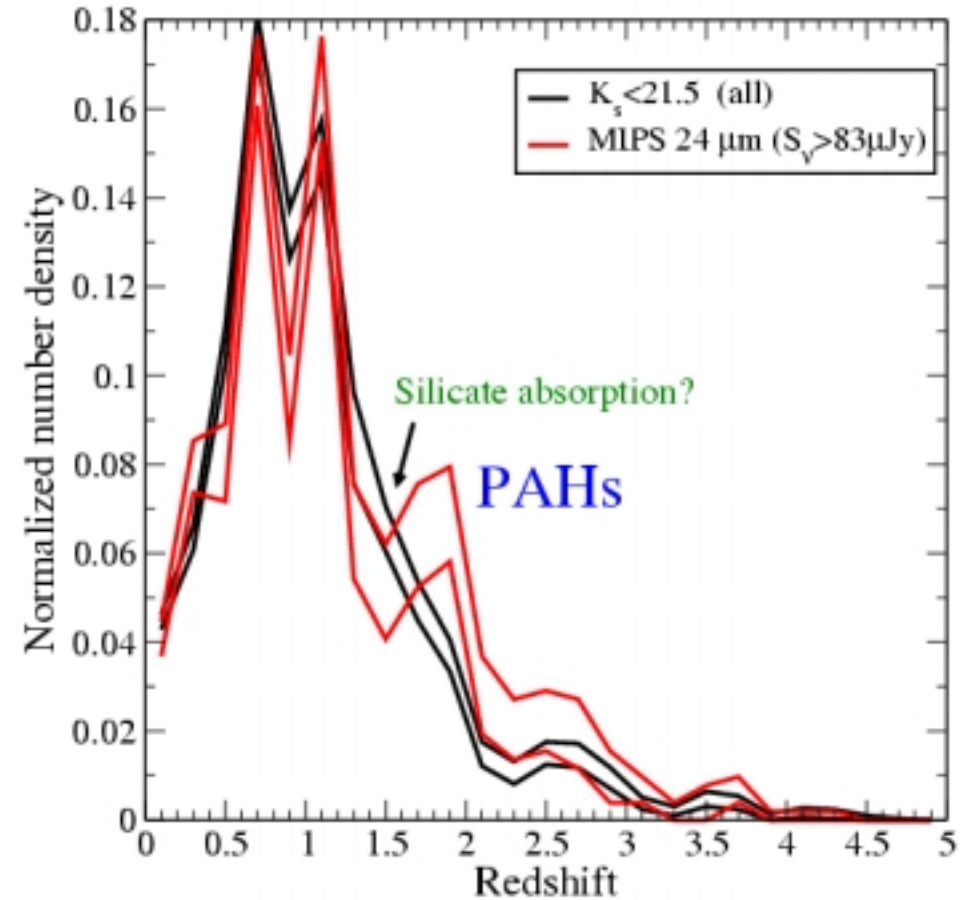


24 μ m Galaxies: e.g. in CDFS



Le Floc'h et al, 2005, ApJ

At $z \sim 1$, $\sim 70\%$ of IR energy density comes from LIRGs

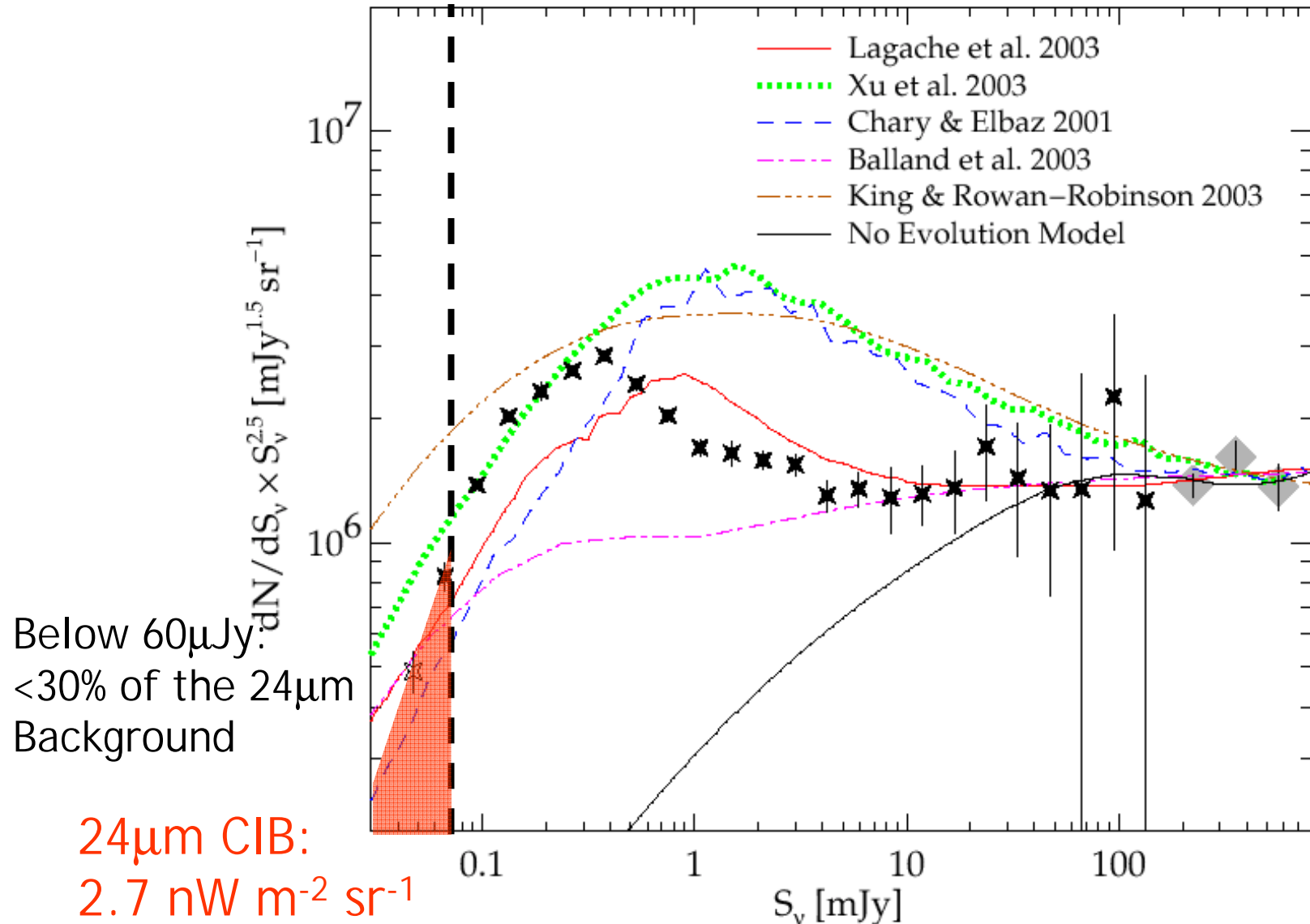


Caputi et al, 2006, ApJ

$\sim 30\%$ of galaxies have $z > 1.5$

+ Luminosity Functions up to $z \sim 2$
(Le Floc'h; Pérez-González; Caputi)

24μm Source Counts & CIB



Papovich et al., 2004

Resolving the FIR CIB

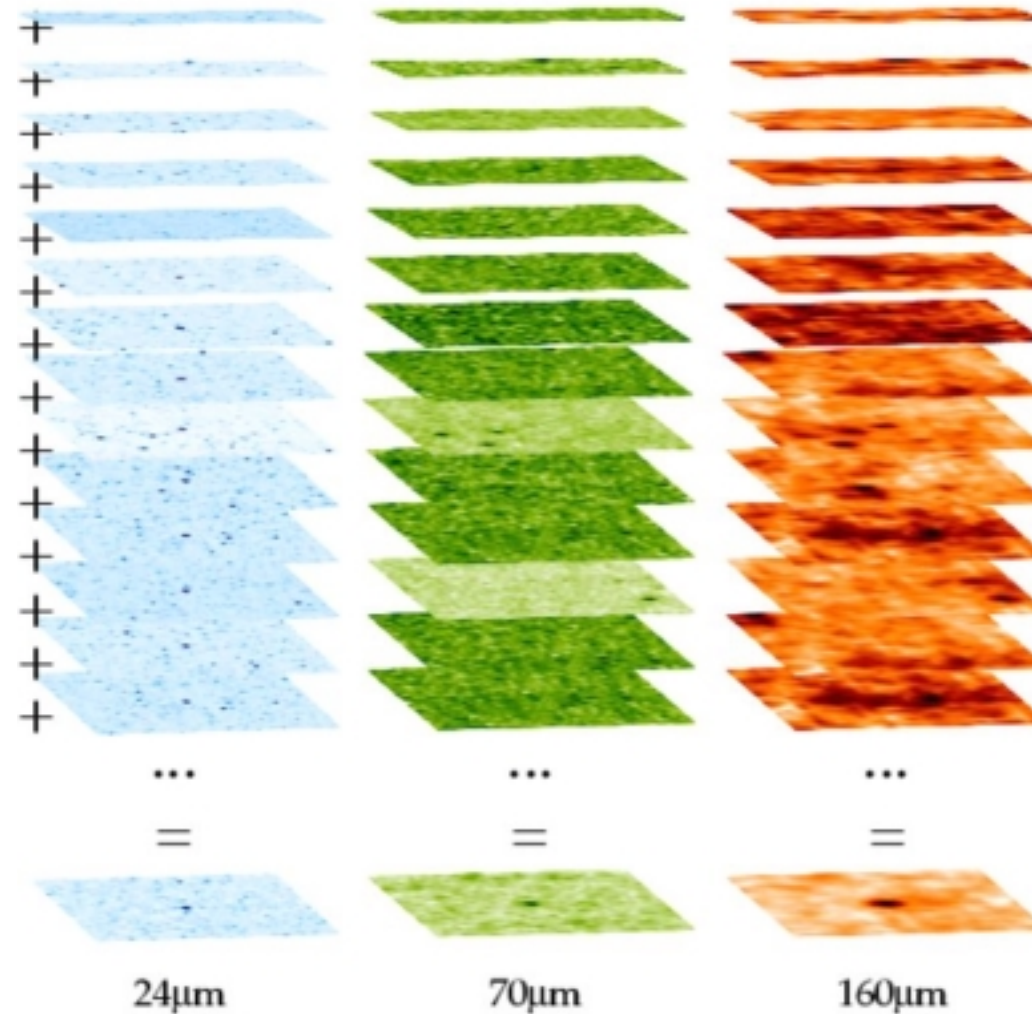


MIPS Stacking Analysis

Dole et al., 2006

Stack ~19000
Galaxies

Watch the
Movie at the
end of this
talk





MIPS 24

Stacking Analysis

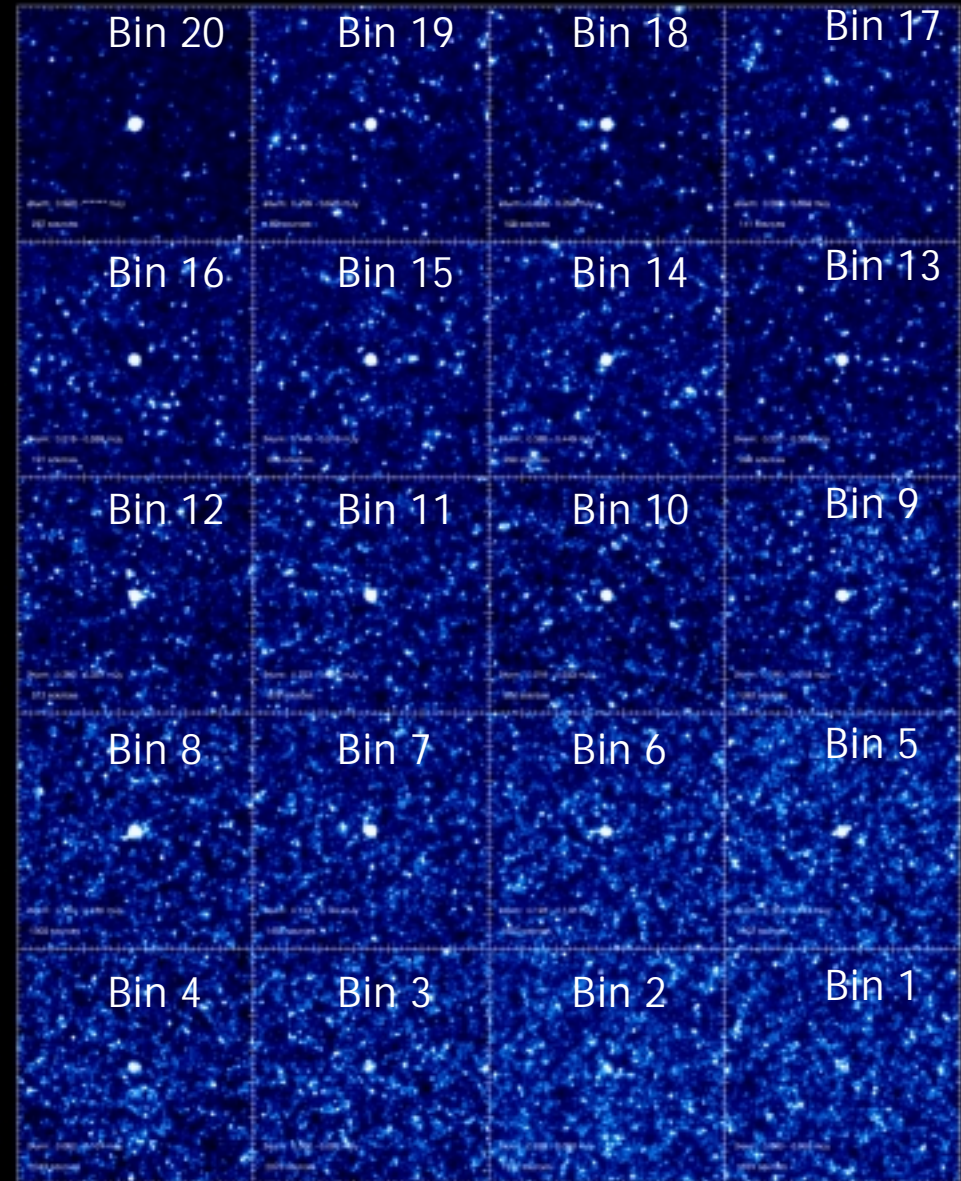
- Use 24 μ m Sources to probe the CIB
- Take Every 24 μ m Source in a Flux Bin
- Add Signal at 24
- Add also the Signal at 70 and 160 at the Same Positions
- Stop at $S_{24} = 60\mu\text{Jy}$ (50% completeness)

3 GTO Fields
~0.9 Sq. Deg.
~19000 sources

Dole et al., 2006, A&A, 451, 417

MIPS 24 μ m

MIPS 24 μ m Stacked Images





MIPS 70

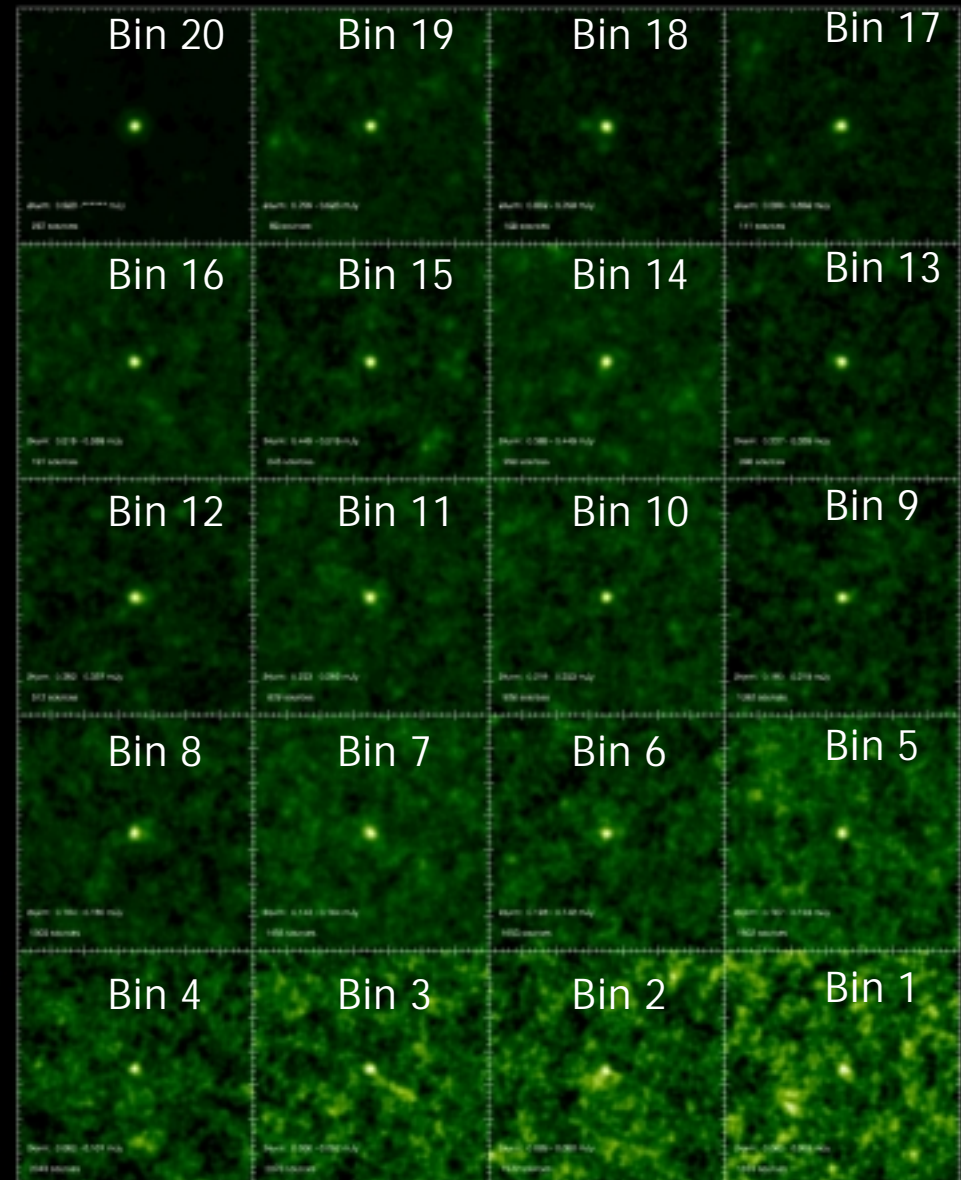
S₂₄ bins:

#	flux range	# added sources
20-	> 0.92 mJy	267 gal
19-	800-920 mJy	83 gal
18-	690-800 mJy	108 gal
17-	590-690 mJy	141 gal
16-	520-590 mJy	191 gal
15-	450-520 mJy	245 gal
14-	390-450 mJy	298 gal
13-	337-390 mJy	396 gal
12-	292-337 mJy	512 gal
11-	250-292 mJy	679 gal
10-	219-250 mJy	820 gal
9-	190-219 mJy	1092 gal
8-	164-190 mJy	1300 gal
7-	142-164 mJy	1458 gal
6-	123-142 mJy	1653 gal
5-	107-123 mJy	1902 gal
4-	92-107 mJy	2040 gal
3-	80-92 mJy	2073 gal
2-	69-80 mJy	1972 gal
1-	60-69 mJy	1851 gal

Dole et al., 2006, A&A, 451, 417

MIPS 70 μm

MIPS 70 μm Stacked Images





MIPS 160

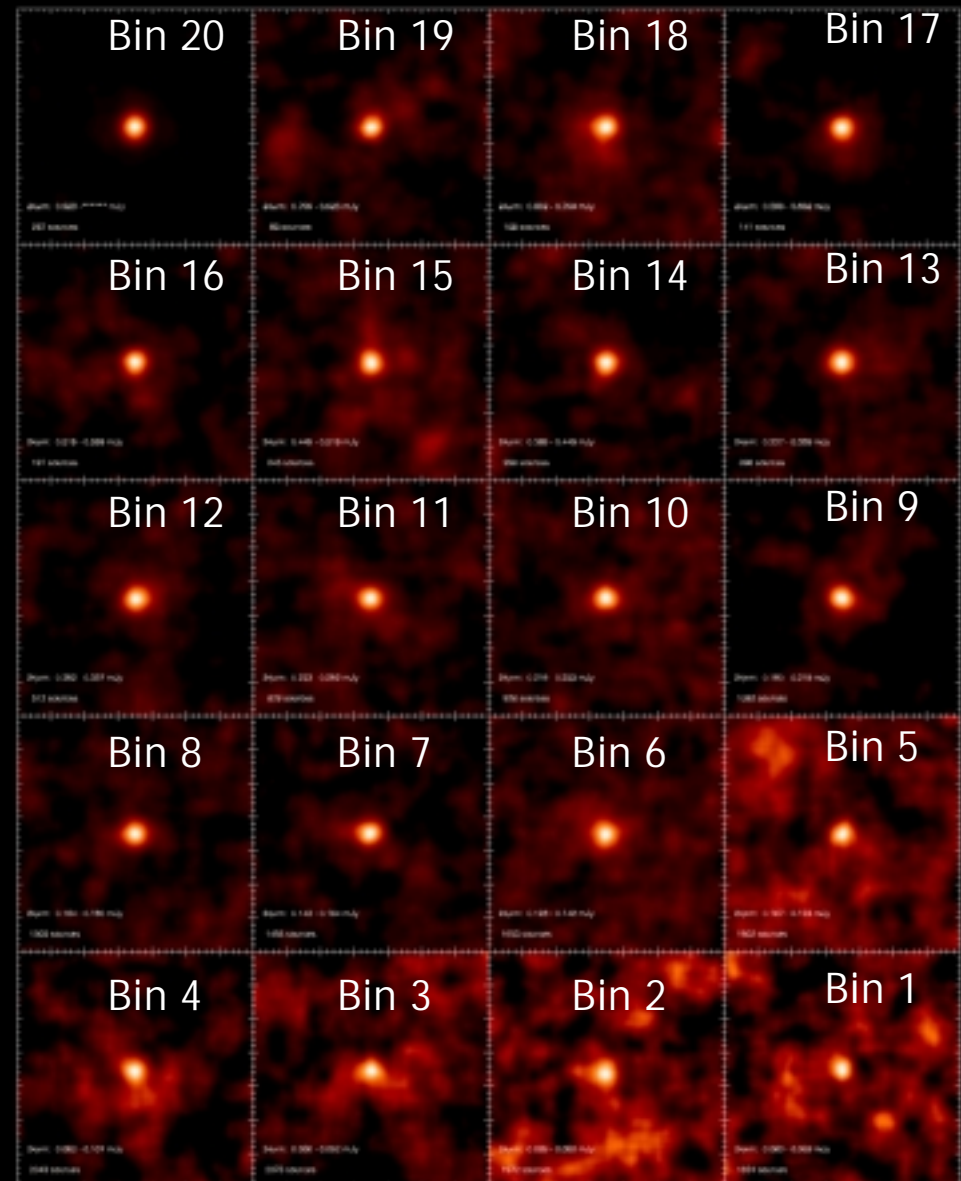
MIPS 160 μm

MIPS 160 μm Stacked Images

S₂₄ bins:

#	flux range	# added sources
20-	> 0.92 mJy	267 gal
19-	800-920 mJy	83 gal
18-	690-800 mJy	108 gal
17-	590-690 mJy	141 gal
16-	520-590 mJy	191 gal
15-	450-520 mJy	245 gal
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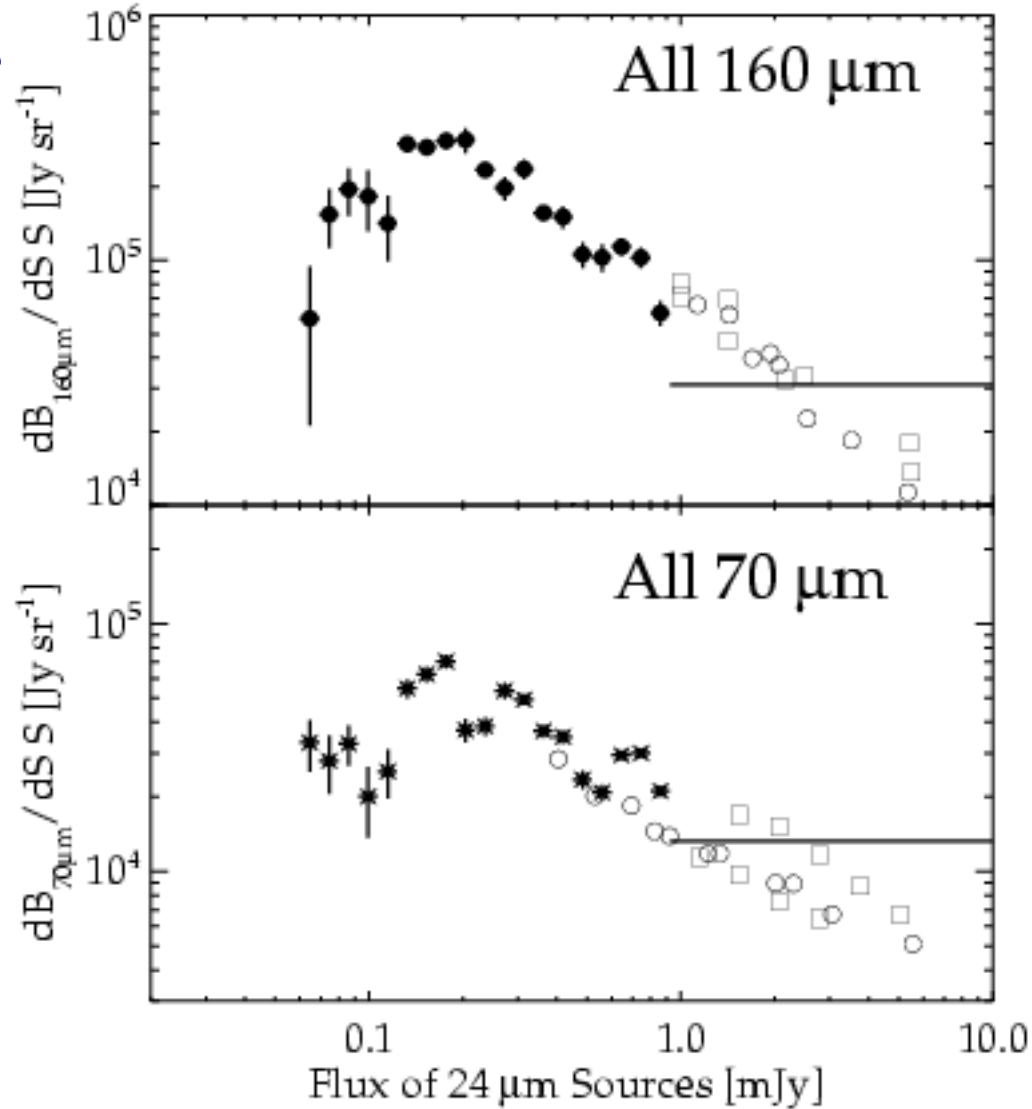
Dole et al., 2006, A&A, 451, 417





CIB Resolved !!!

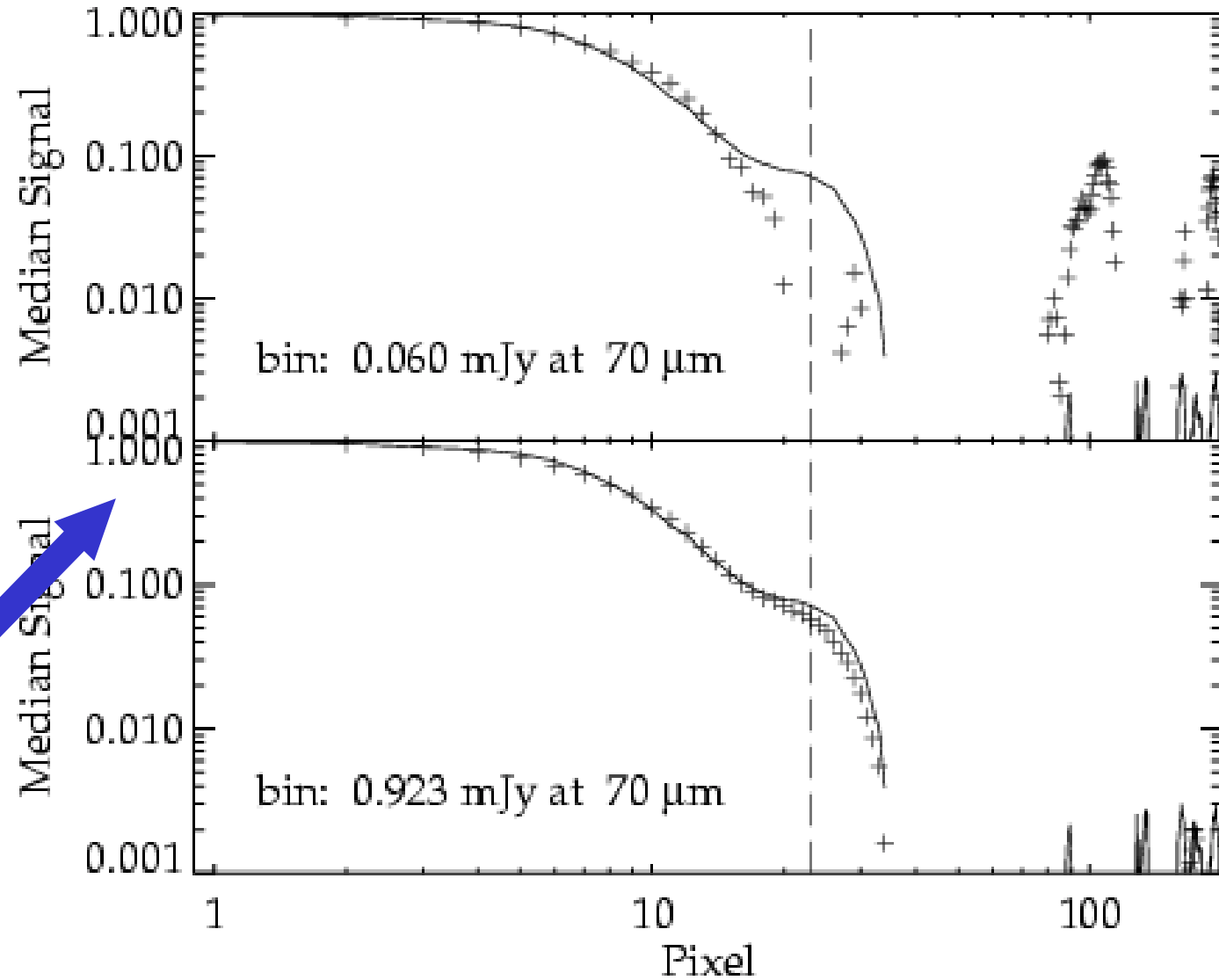
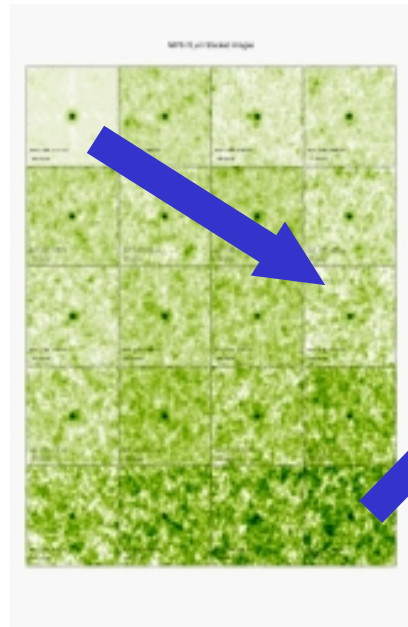
Contributions to the CIB



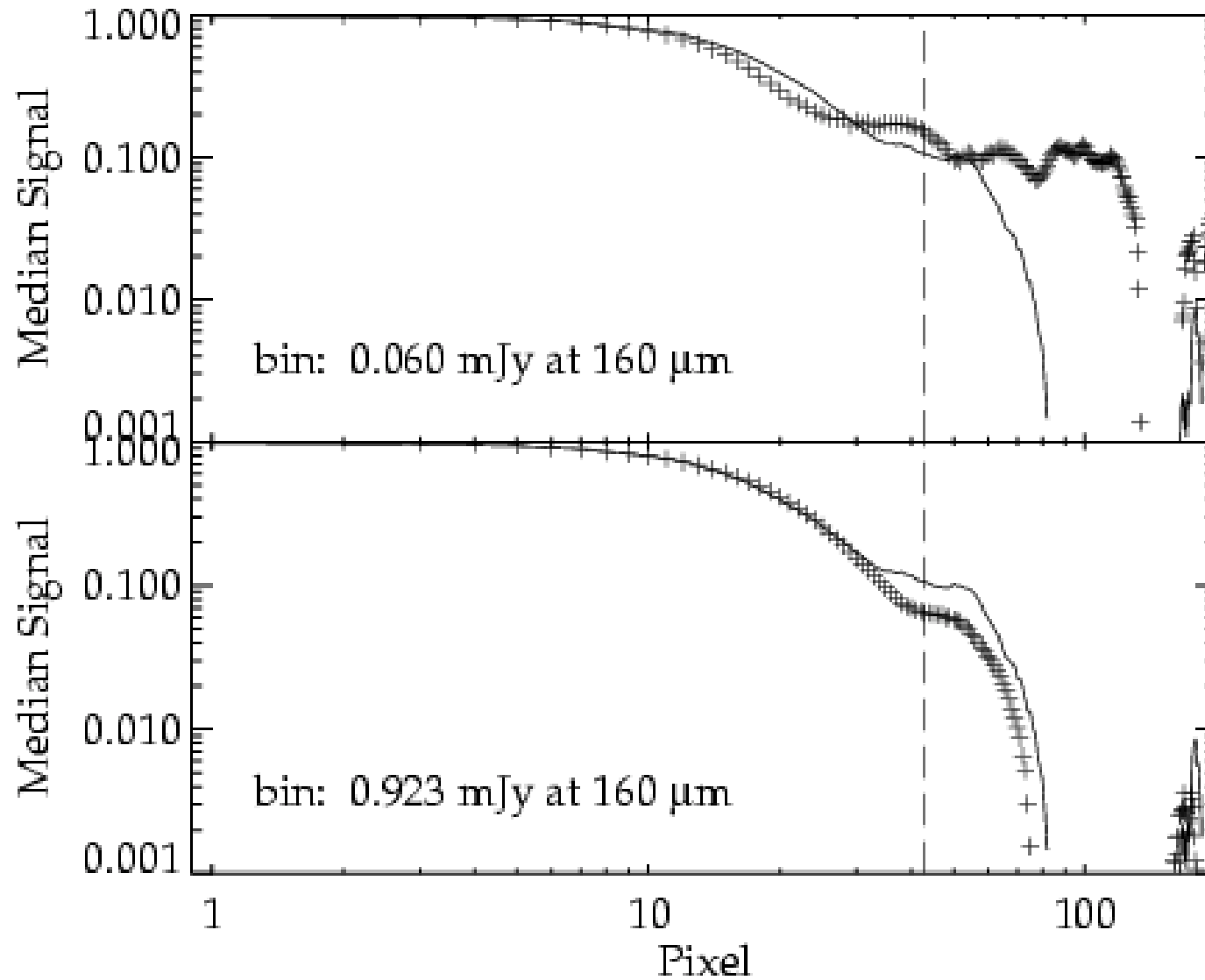
Source Counts from:
-Dole et al, 2004
-Frayser et al, 2005

CIB resolved at:
- 1.6mJy @70
- 6mJy @160

Radial Profiles at 70 μ m

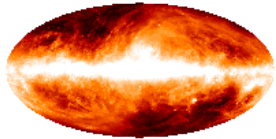


Radial Profiles at 160 μ m



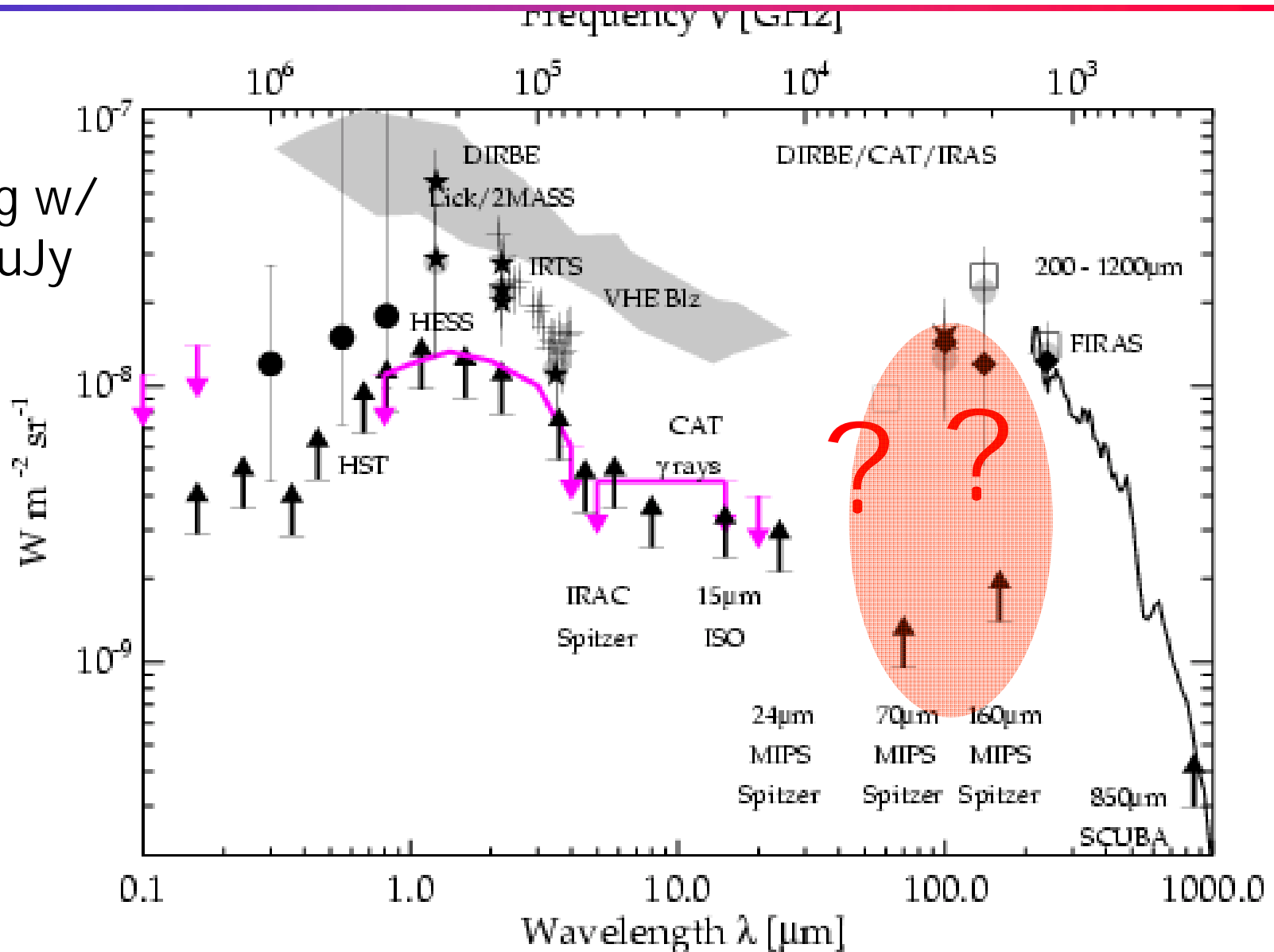
Results on the CIB

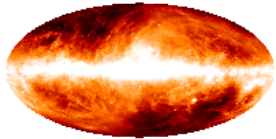




EBL and CIB

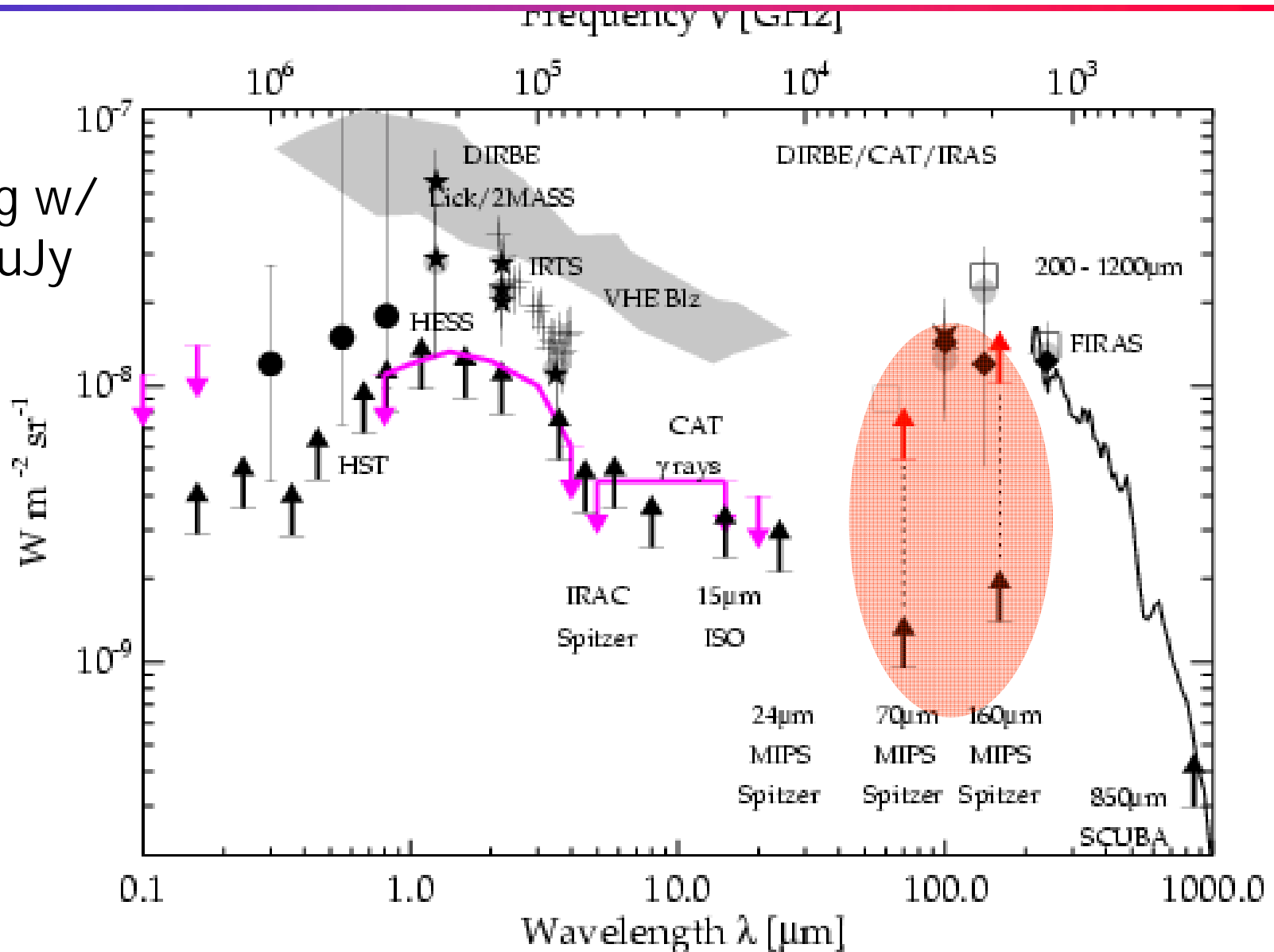
Stacking w/
 $S_{24} > 60 \mu\text{Jy}$

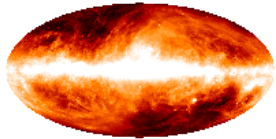




EBL and CIB

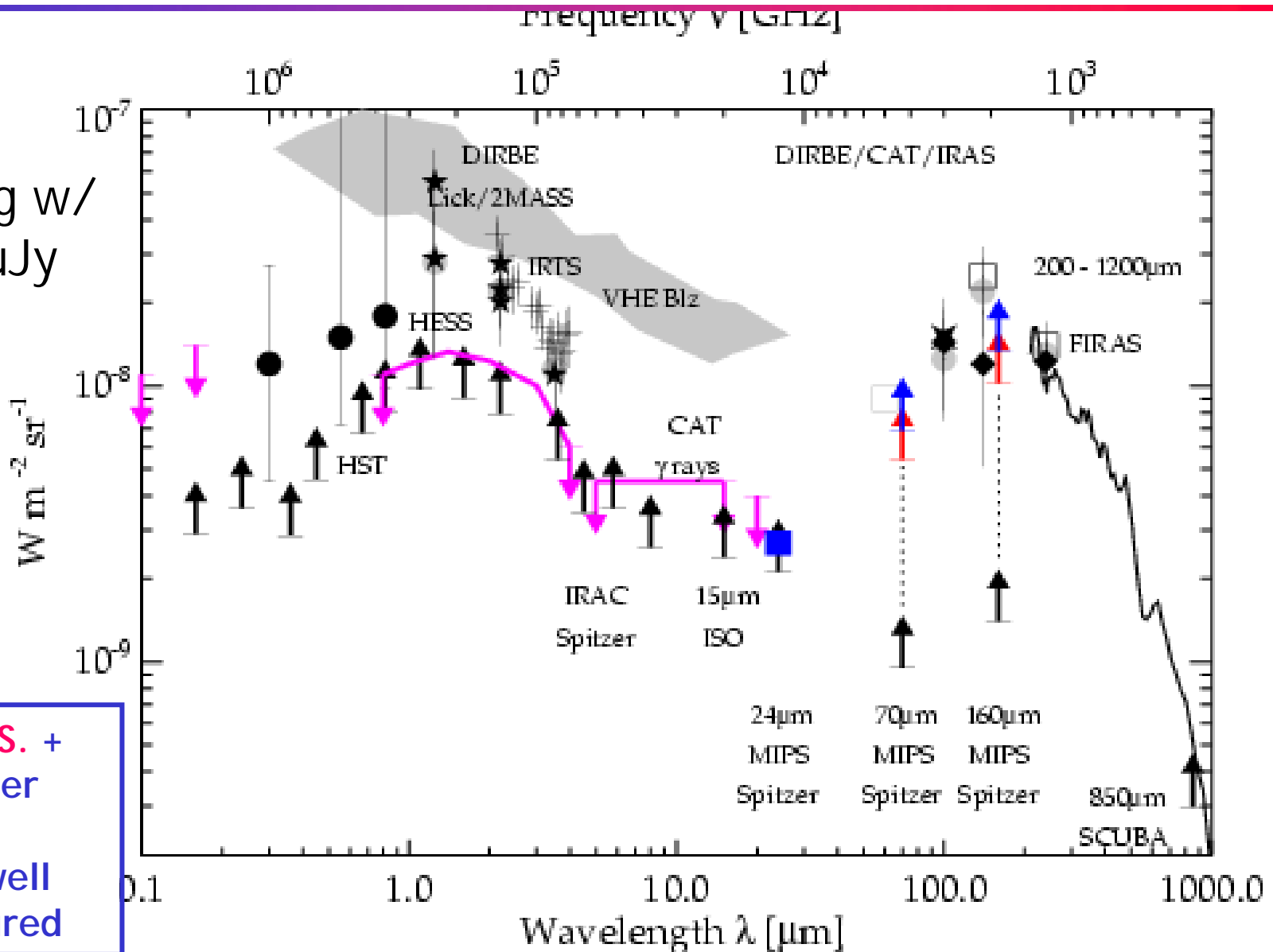
Stacking w/
 $S_{24} > 60 \mu\text{Jy}$





EBL and CIB

Stacking w/
 $S_{24} > 0.0 \mu\text{Jy}$





Cosmic Infrared Background

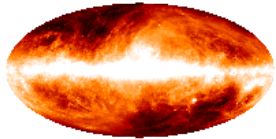
Galaxies Making-up the CIB

- 24um galaxies [$S_{24} > 60 \mu\text{Jy}$] contribute to ~80% of the FIR CIB
 - Measured, completely model-independent
 - Confirms Elbaz et al (2002) model-dependent result
- MIR galaxies are thus **good tracers** of galaxies making-up the bulk of the CIB
- We can also probe the CIB deeper for the $S_{24} < 60 \mu\text{Jy}$ galaxies ...

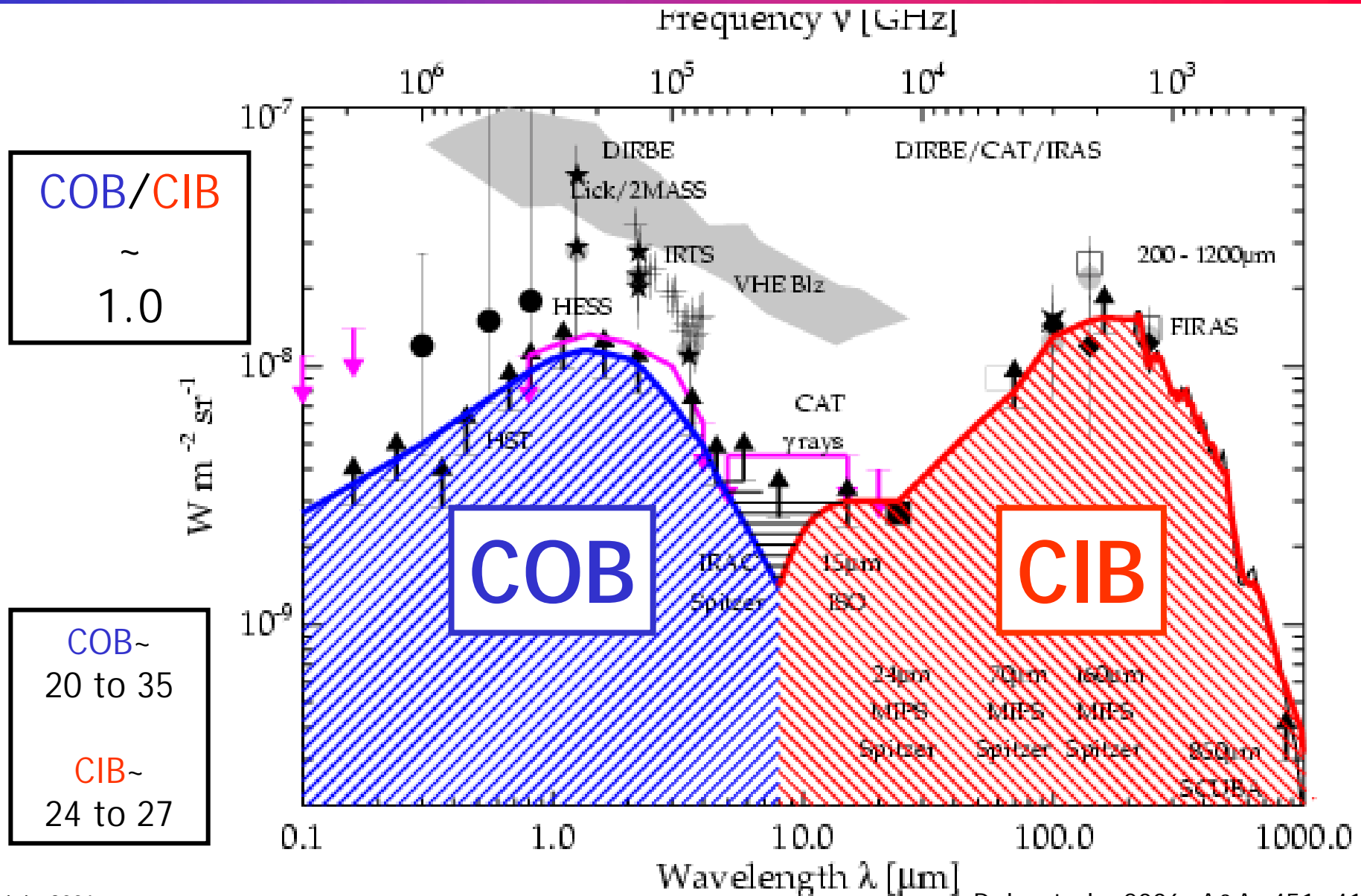
New Estimate of the FIR CIB

- Using Stacking Analysis [$S_{24} > 60 \mu\text{Jy}$]
- Using unresolved bkg at 24um [$S_{24} < 60 \mu\text{Jy}$]
- Using 70/24 and 160/24 observed colors

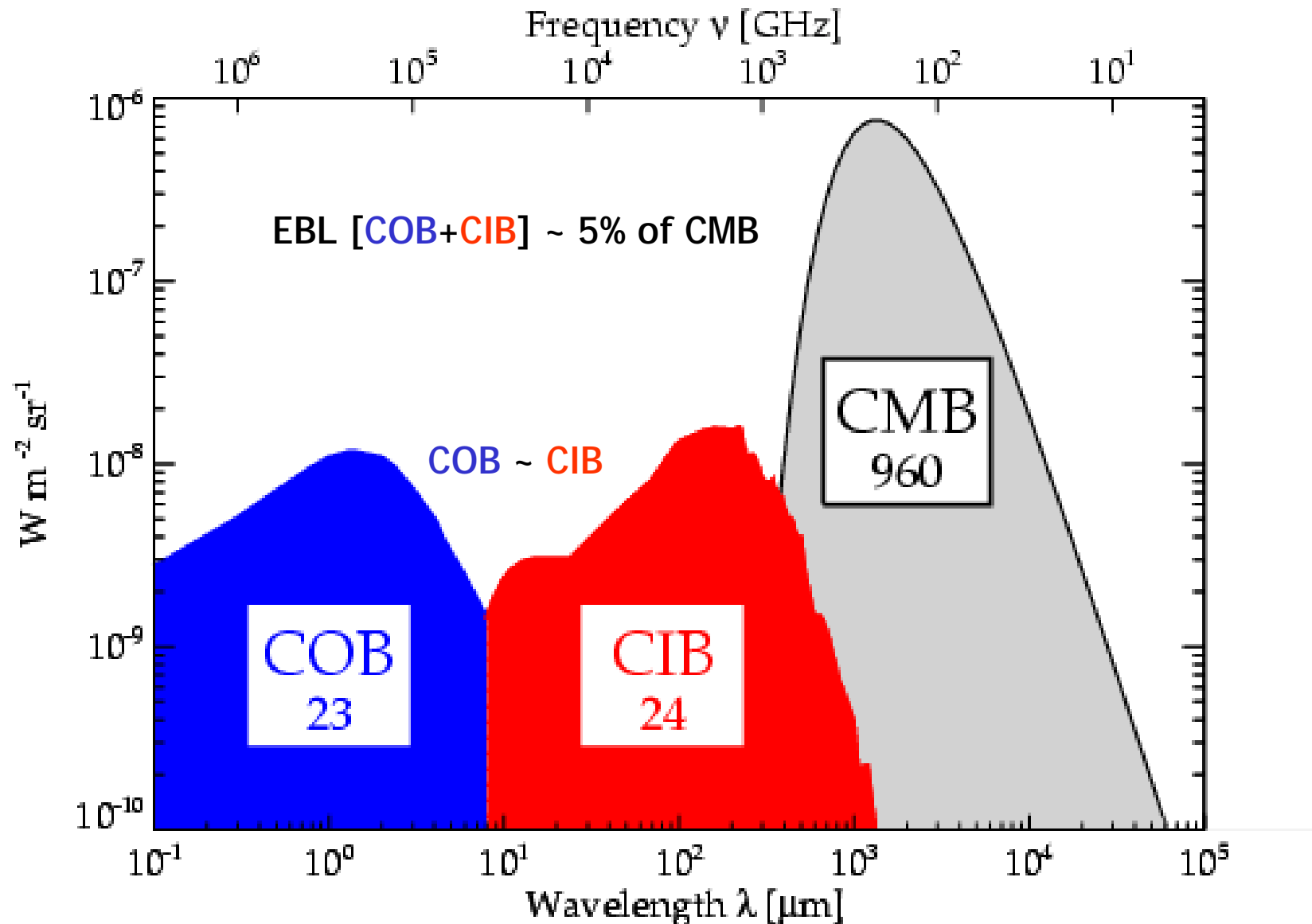
CIB	24 um	70 um	160 um
CIB nW m ⁻² sr ⁻¹	2.7	7.1 +/- 1.0	13.4 +/- 1.7
CIB MJy/sr	0.021	0.16 +/- 0.02	0.71 +/- 0.09



Extragalactic Background Light



Universe' Spectral Energy Distribution



Conclusions

○ Spitzer Resolves the CIB

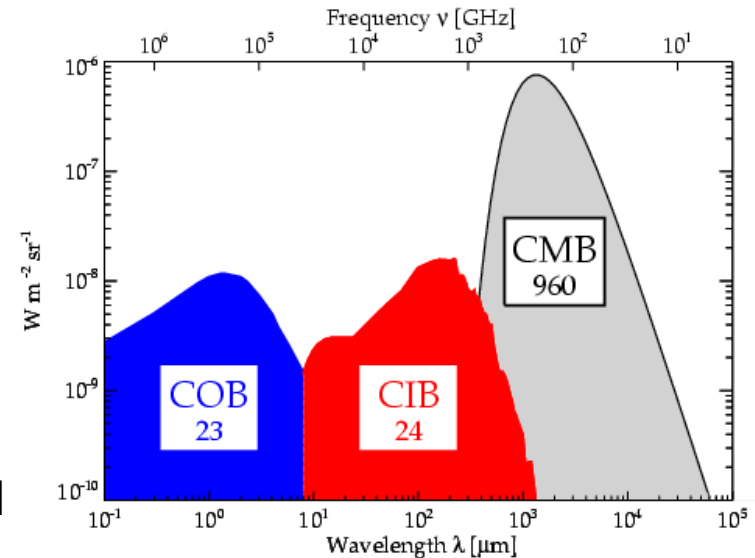
- ❑ FIR CIB now resolved up to 80%; new CIB estimates
- ❑ MIR traces well the FIR (CIB peak) population

○ CIB vs COB vs CMB

- ❑ Extragalactic Background Light SED now **well constrained**
- ❑ Budget for Galaxy Formation & Evolution
- ❑ COB ~ CIB ~ 24/23 nW m⁻² sr⁻¹
- ❑ EBL: for 1 visible photon => 115 IR photons
- ❑ EBL ~ 5% CMB

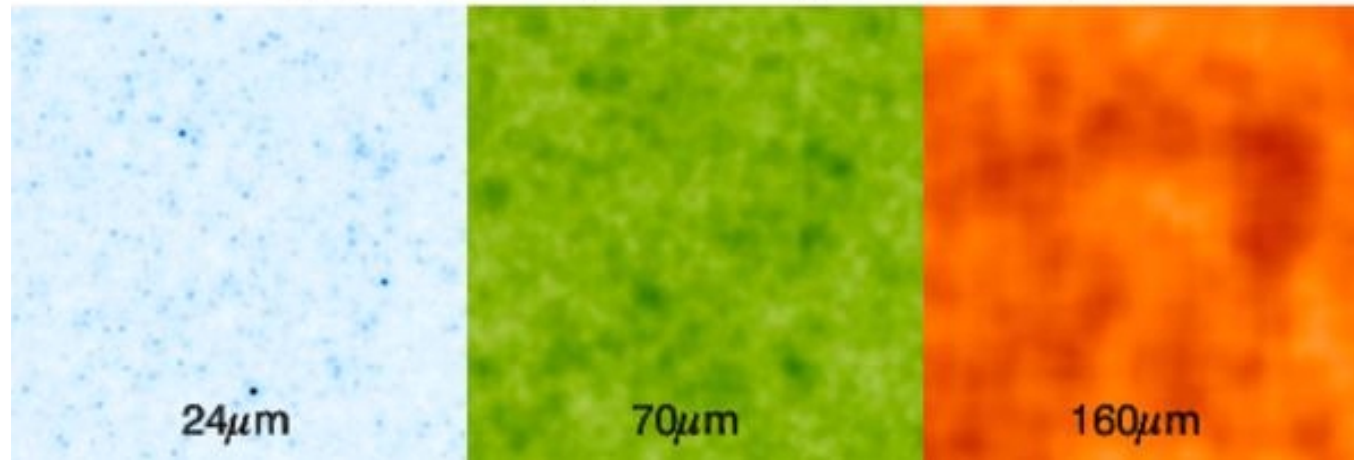
○ Sources of the CIB

- ❑ Bulk of the CIB: z ~1 to 2
- ❑ FIR SEDs of faint sources are being measured
 - ❑ up to z~1.5
- ❑ Next/Ongoing: Deep IRS spectroscopy + Deep Imaging Spitzer programs
- ❑ **Stacking Analysis**: very promising



Movie of the Stacking Analysis

Stacking Analysis Technique Check
Example of 300 stacked images at random sky positions. Dole et al. (2006)



Grab it here:

<http://www.ias.u-psud.fr/irgalaxies>

<http://insu.cnrs.fr/web/article/art.php?art=1747>

<http://www.spitzer.caltech.edu/Media/happenings/20060420/>



<http://www.ias.u-psud.fr/irgalaxies>

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Emeric Le Floc'h
Eiichi Egami
Pablo Pérez-González

+ MIPS People
+ SSC, IRAC, IRS people
+ Legacy Teams (SWIRE, GOODS)