



What do global p -modes tell us about large scale solar flows?

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What this work is about

- (Re)calculation of effects of Zonal (rotation) and poloidal solar flows (meridional circulation , giant cells) on p -mode splitting coefficients. *Leads to much lower estimates for frequency shifts!*

Previous related work:

Roth, Howe and Komm, 2002, A&A, 396, 243

Roth and Stix (1999, 2003, 2008)

- Compare theoretical splitting coefficients with GONG as well as MDI data sets and *put upper limits on the magnitudes of these flows.*



The technique

- Quasi-degenerate Perturbation Theory (Lavelly & Ritzwoller, 1992)

Couples two modes with slightly different un-perturbed frequencies (for us $|\omega_2 - \omega_1| < 100 \mu\text{Hz}$)

$$\Delta = \omega_2^2 - \omega_1^2$$

$$\begin{bmatrix} H_{11} - \Delta & H_{12} \\ H_{21} & H_{22} \end{bmatrix} \rightarrow \lambda = H_{22} - \frac{|H_{12}|^2}{H_{11} - H_{22} - \Delta}$$

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In contrast *degenerate perturbation theory* (DPT) couples only exactly degenerate unperturbed modes



The Model

- Standard solar model with OPAL equation of state, OPAL opacities, convective flux calculation (Canuto & Mazzitelli, 1991)
- Eigenfunctions and frequencies from a *Solar pulsation code* (with non-adiabatic effects and without Cowling's approx, H. M. Antia, 2002, ver 2.2)

QDPT: Application to rotation

$$\Omega r \sin \theta = -w_1^0 \partial_\theta Y_1^0 - w_3^0 \partial_\theta Y_3^0$$

$$\omega_{nlm} = \omega_{nl} + \sum_q a_q^{(nl)} P_q^l(m)$$

- Zonal flows affects p -modes in two ways:

1. Coriolis force, linear in Ω (odd splitting coefficients a_{2q+1} from DPT)

N-S asymmetric Ω gives even splitting coefficients a_{2q} :

with $w_2^0 = -7.8 \pm 0.3 \text{ ms}^{-1}$ (Hathaway et al 1996), we calculate $a_2 \sim 0.1 \text{ nHz}$

2. Centrifugal force, Ω^2 (DPT gives even splitting coeff, a_{2q} Antia et al. 2000: $a_2 < 25 \text{ nHz/l}$)



QDPT: Application to rotation

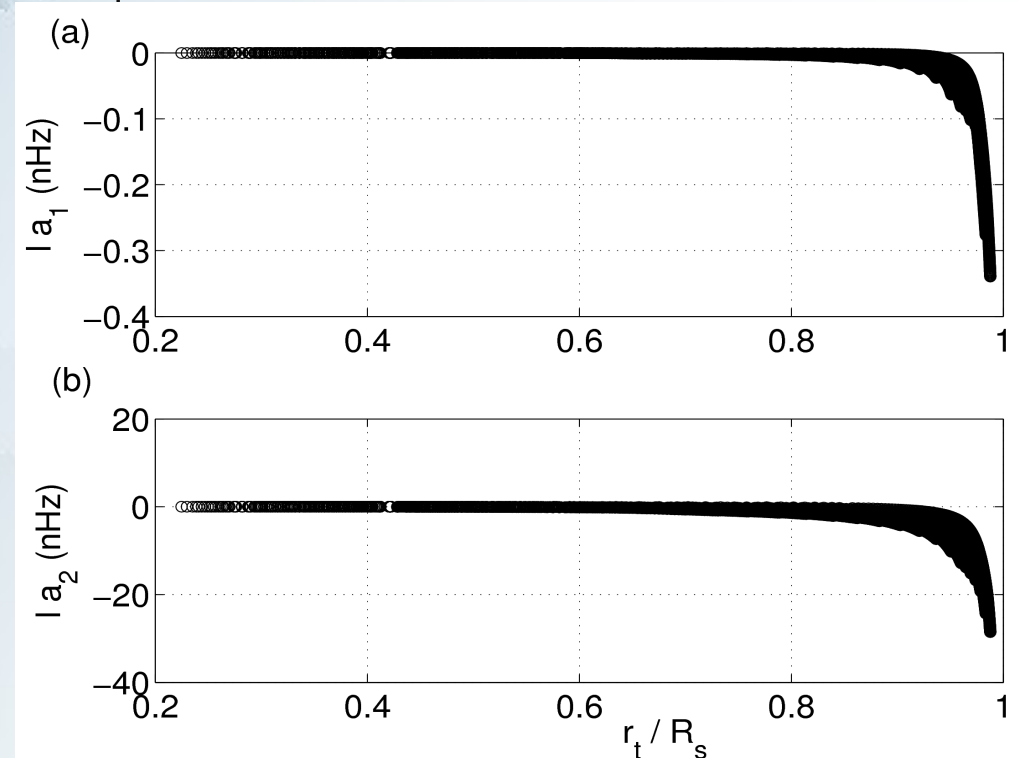
QDPT can calculate effect of Ω (Coriolis force) on a_{2q} as a *second order correction*

$$\omega_{nlm} = \left[\omega_{nl}^2 + \underbrace{H_{22}}_{m^2} - \frac{\underbrace{|H_{12}|^2}_{m^4}}{H_{11} - H_{22} - \Delta} \right]^{1/2}; H_{22} \approx 450 \text{ m nHz} \gg H_{21}$$

For w_1^0 term $H_{12} = 0!$

For $w_3^0 = 34.9 \text{ ms}^{-1}$ $r > 0.7$,

$la_1 \sim 0.4 \text{ nHz}$; $la_2 < 35 \text{ nHz}$

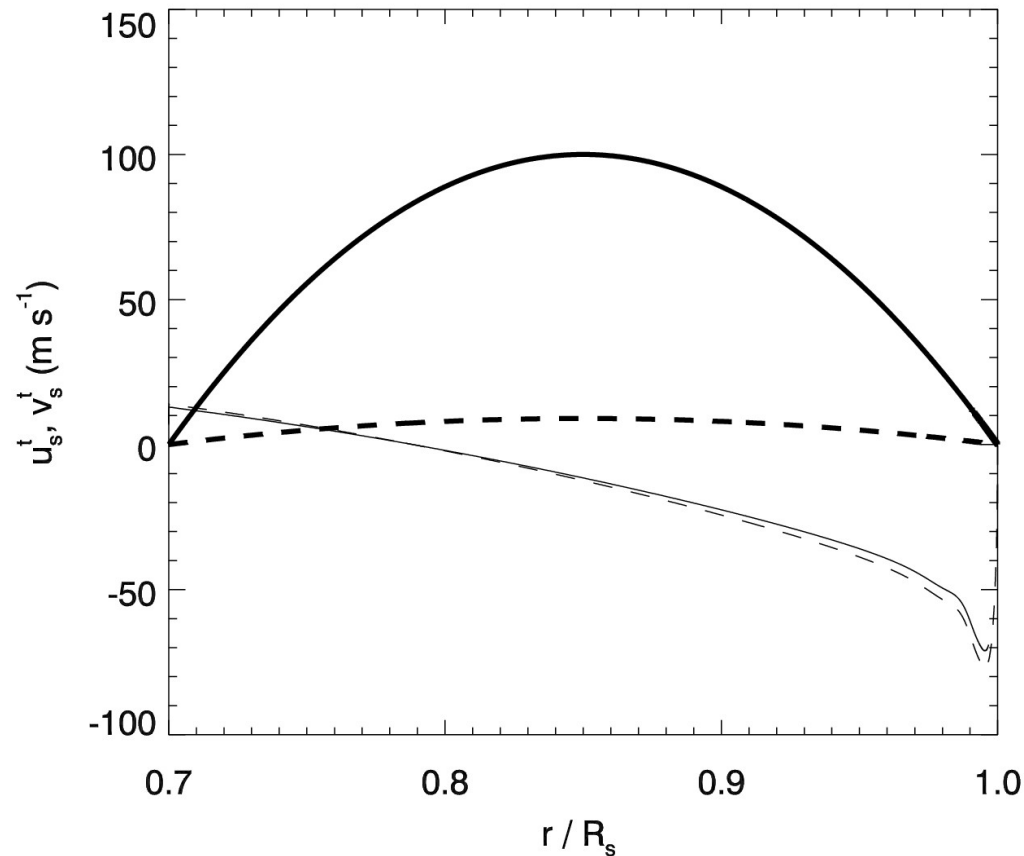


QDPT: Non Zonal flows

$$(u_0 = 9 \text{ ms}^{-1}, s = 2, t=0, Y_2^0)$$

- Meridional circulation

$$\mathbf{v}(r, \theta, \varphi) = u_s^t(r) Y_s^t + v_s^t(r) \partial_\theta Y_s^t$$



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$$\omega_{nlm} = \left[\omega_{nl}^2 + \underbrace{H_{22}}_0 - \frac{|H_{12}|^2}{H_{11} - H_{22} - \Delta} \right]^{1/2} \mathbf{v}^2$$

QDPT is necessary to provide the first correction due to poloidal flows

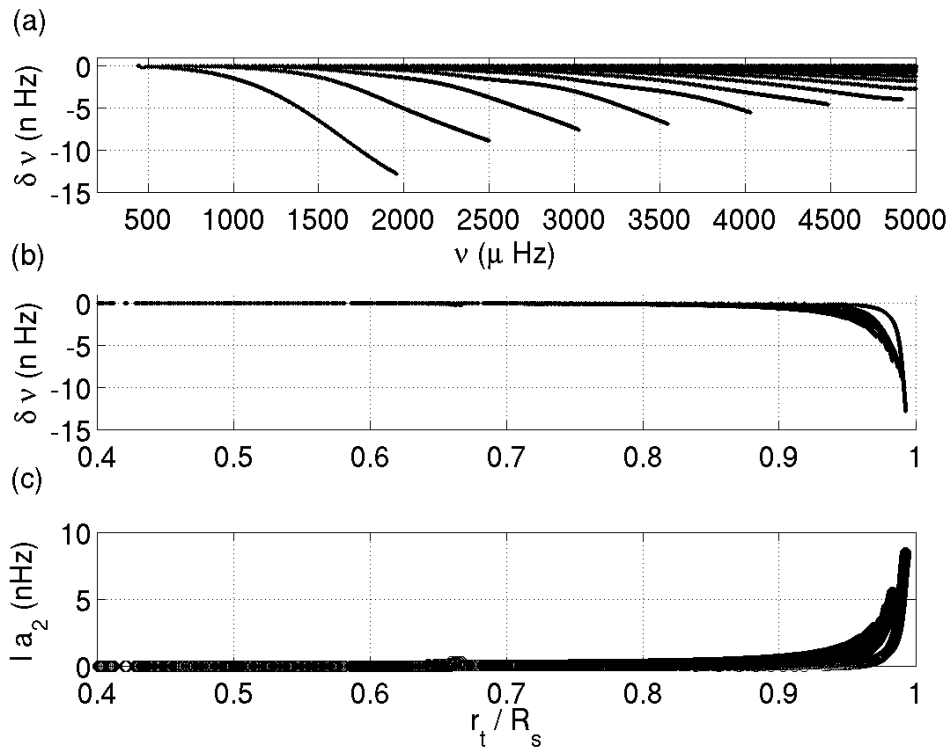
$$\delta\omega_{nlm} \approx \frac{H_{12}^2}{2\omega_{nl} \Delta}$$

Where, Δ = difference in squared frequency of the coupling modes



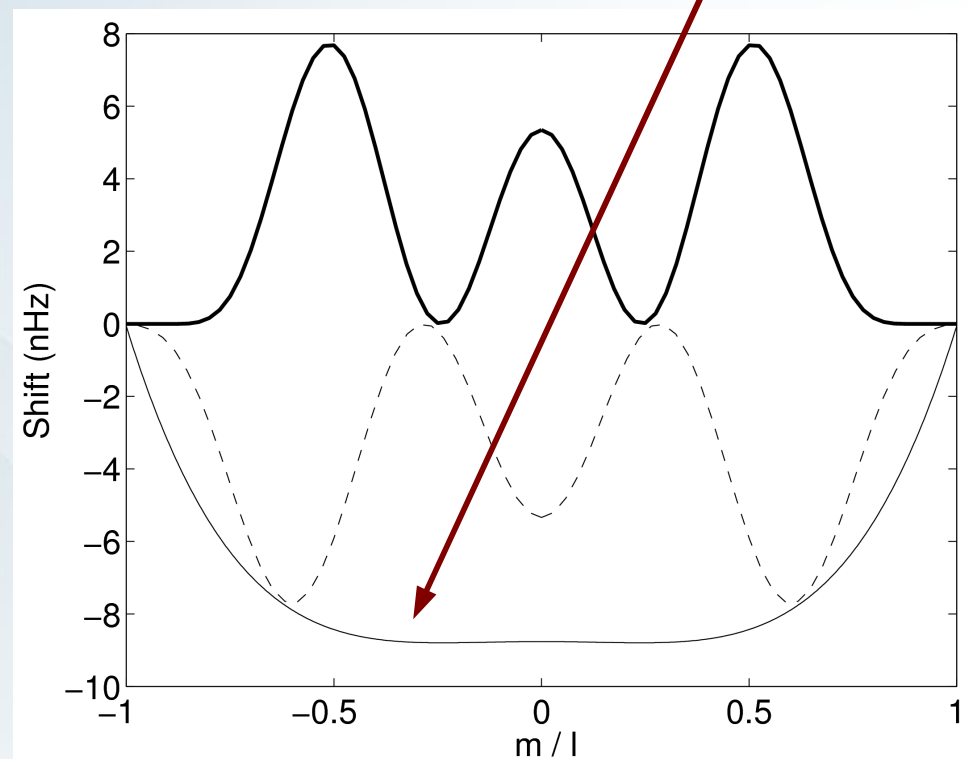
QDPT: Meridional flow

$$(u_0 = 9 \text{ ms}^{-1}, s = 2, t=0, Y_2^0)$$



$$\delta\omega(m) \propto m^{2q}, u_0^2$$

$$\text{For } (n, l) = (1, 292)$$



la_1 and la_2 as a function of
turning point radius r_t

Max $\delta\omega \sim -12.5$ nHz!

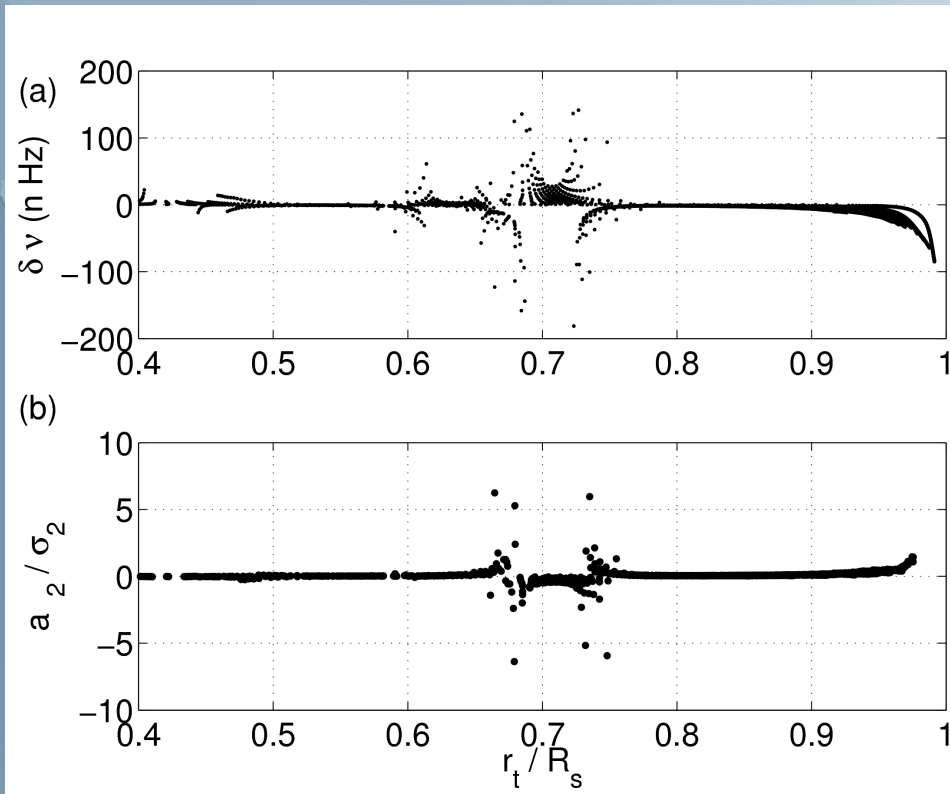
Max $la_2 \sim 8.5$ nHz!

Chatterjee, P. & Antia, H. M.



QDPT: Meridional flow

$$(u_0 = 100 \text{ ms}^{-1}, s = 8, t=0, Y_8^0)$$



Increasing s allows more multiplets with $|\omega_1 - \omega_2| < 100 \mu\text{Hz}$ to couple.

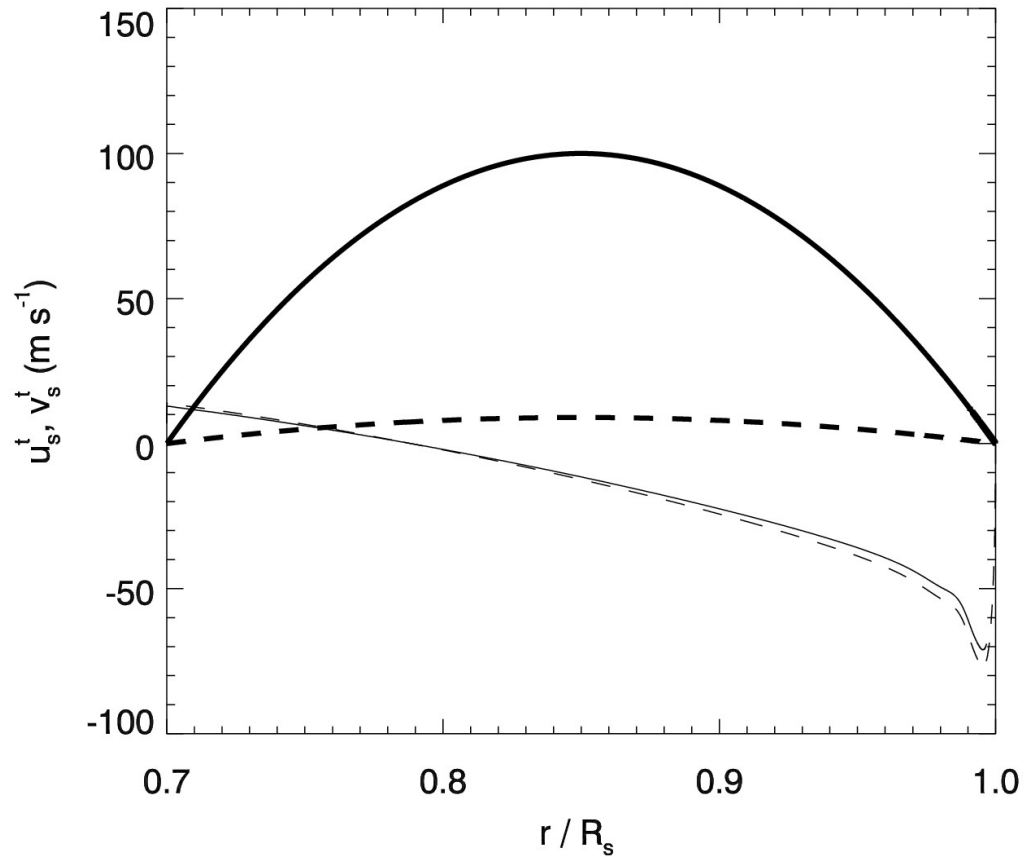
More instances of “near degeneracy” such that $\delta v \propto \Delta^{-1}$ can become very large

where, Δ = difference in squared frequency of the coupling modes

Example: $(n, l) = (17, 56)$ and $(16, 64)$ have $\Delta/2\nu = -0.06\mu\text{Hz}$
 $la_2 = 480 \text{ nHz}$

QDPT: Giant Cells

($u_0 = 100 \text{ ms}^{-1}$, $s = 8$, $t=4, 8, Y_8^8 \ Y_8^4$)



QDPT: Giant Cells

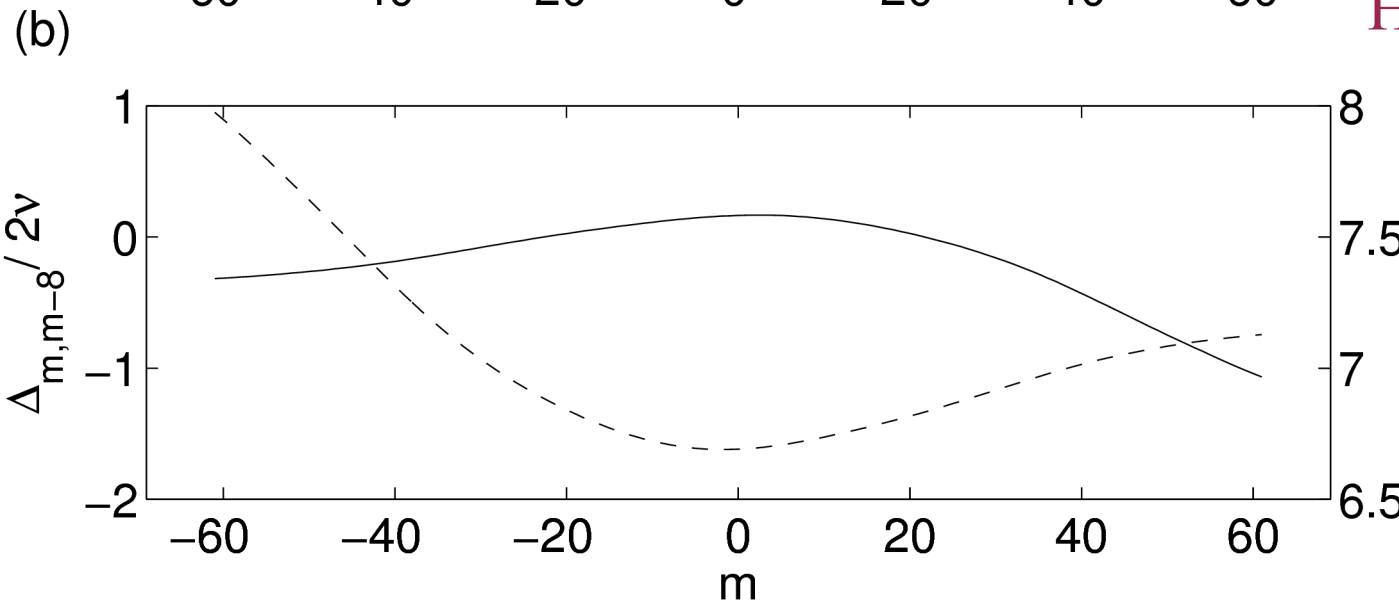
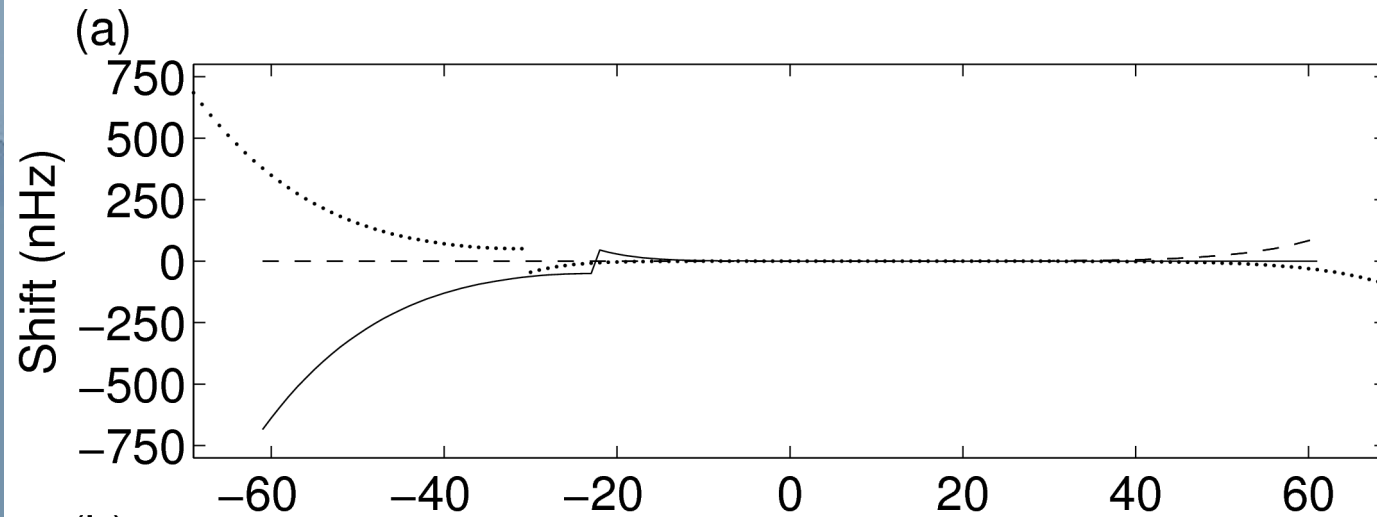
$$(u_0 = 100 \text{ ms}^{-1}, s = 8, t=4, 8, Y_8^8 \quad Y_8^4)$$

- Expression for $u(r, \theta, \varphi)$ involves $Y_8^4(\theta, \varphi)$ or $Y_8^8(\theta, \varphi)$ (banana cells) \Rightarrow coupling between different m, m' of p -modes.
- Important to take effect of rotational splitting on Δ , before calculating the effect of flows with $t \neq 0$.
- Basically $\Delta \equiv \Delta(m)$ unlike for $t = 0$ cases.
- Rotational splittings obtained from temporally averaged GONG data.
- Giant cells cause splittings asymmetric about $m = 0$; odd a_{2q+1} affected. To what extent is rotational inversion affected?



QDPT: Giant Cells

$$(u_0 = 100 \text{ ms}^{-1}, s = 8, t=8, Y_8^8)$$

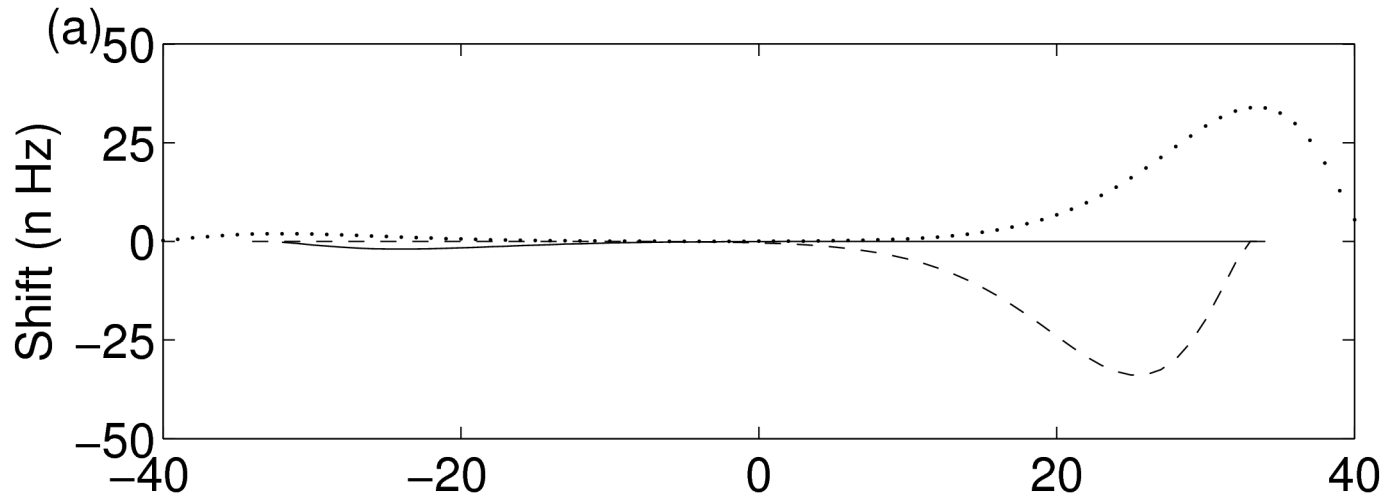


$\Delta / 2v$ (μHz) due to
 n, l n, l
 $(18, 61, m) \Leftrightarrow (17, 69, m \pm 8)$
Has a “near degeneracy”
at $m = -22$
 $\Delta(m) = 0$



QDPT: Giant Cells

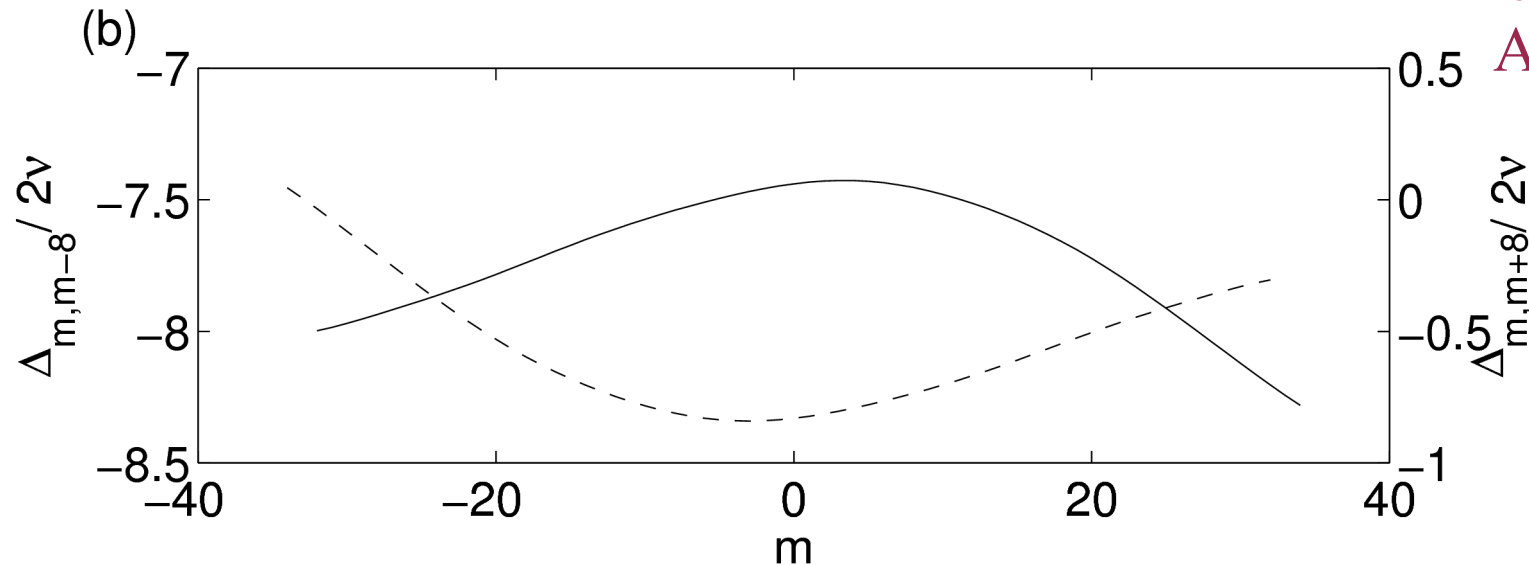
($u_0 = 100 \text{ ms}^{-1}$, $s = 8$, $t=8$, Banana cells)



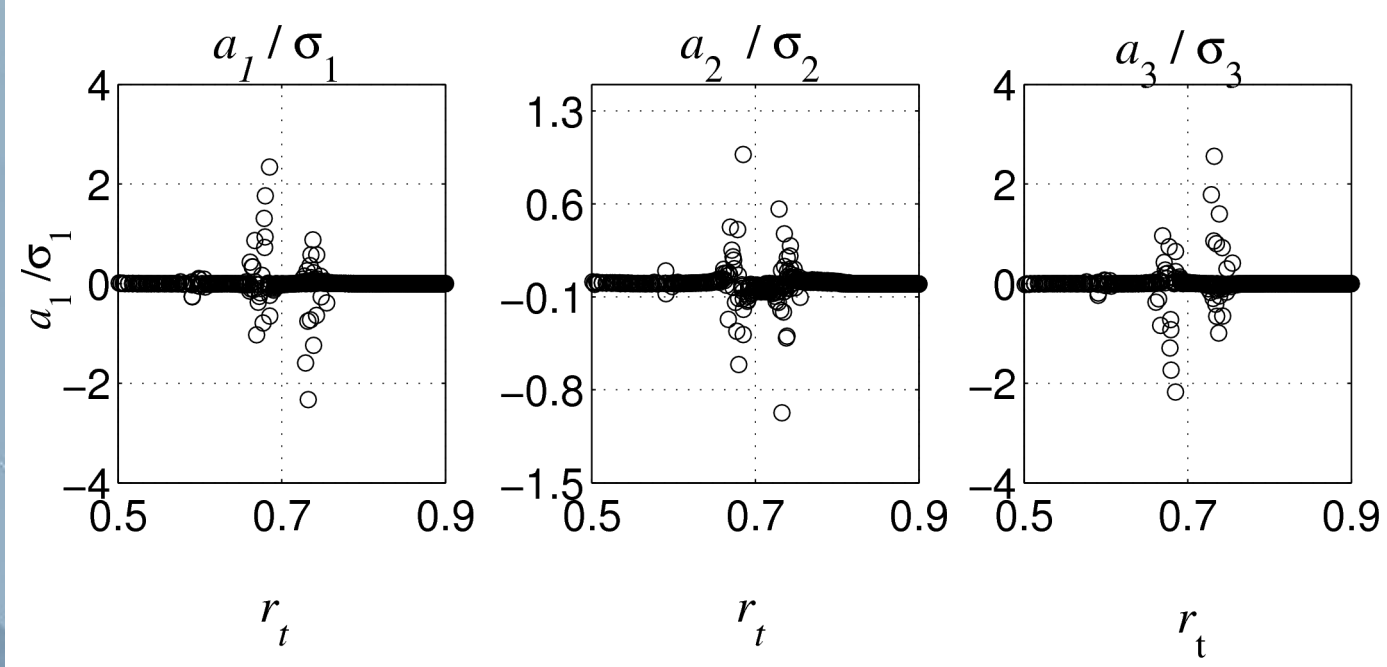
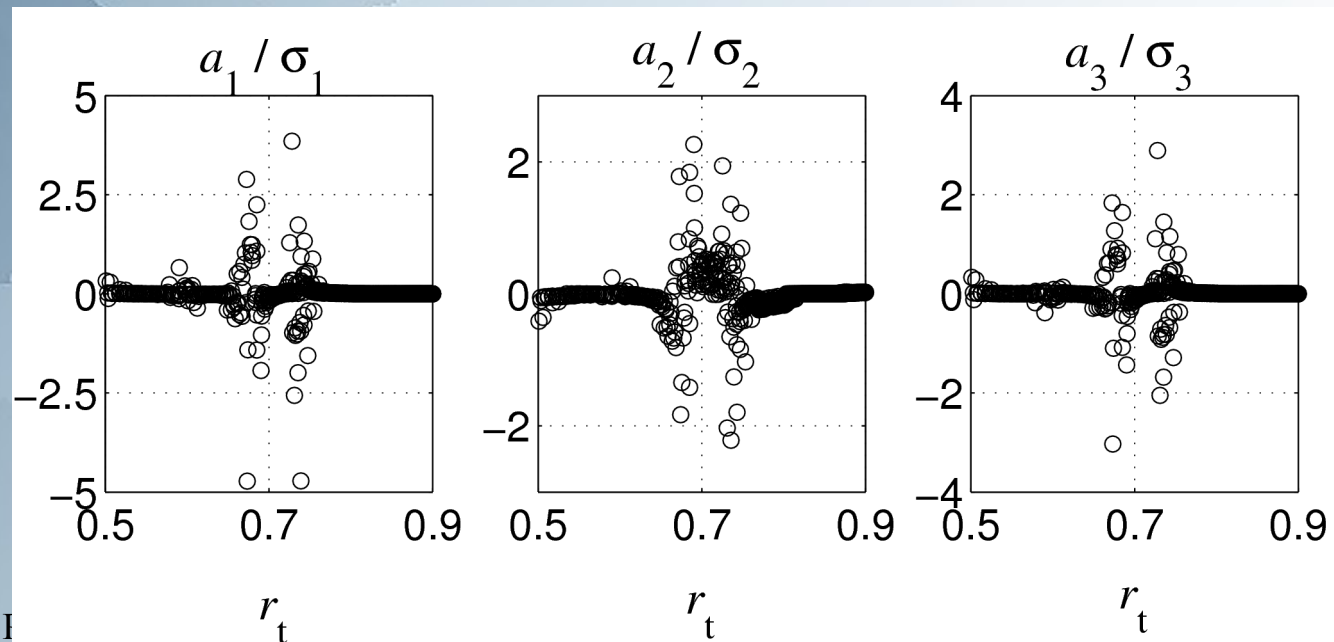
$\Delta / 2v$ (μHz) due to
 n, l n, l
(15, 34, m) \leftrightarrow (14, 40, $m \pm 8$)

Asymmetric about $m=0$

Affects odd splitting
coeff.

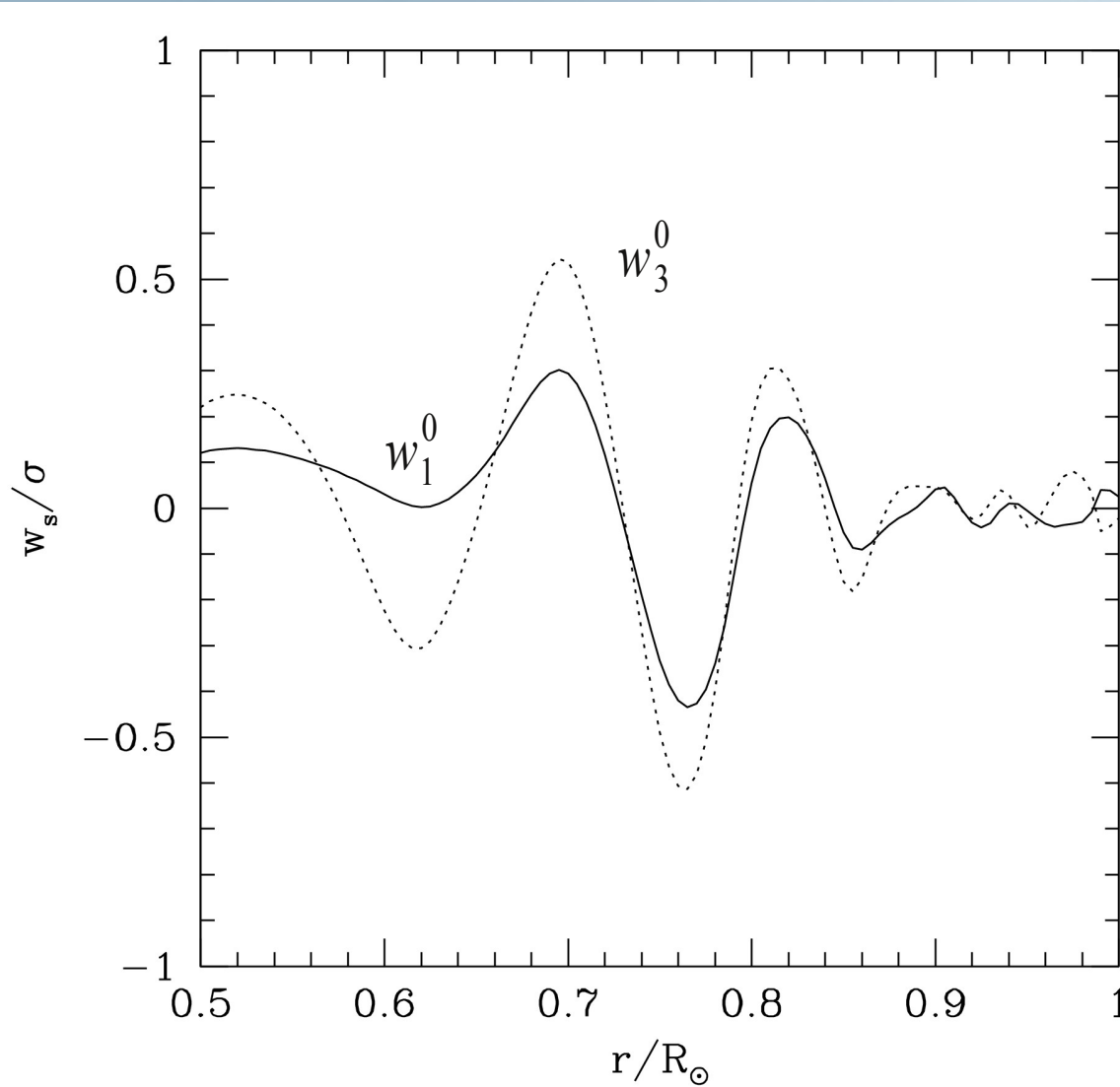


QDPT: Giant Cells


 $Y_8^4(\theta, \varphi)$

 $Y_8^8(\theta, \varphi)$
 (Banana
 cells)



QDPT: 1.5d inversion



To what extent is rotational inversion distorted?
Ans: No appreciable distortion

For Y_8^8 flow,
1.5d helioseismic inversion
using errors in GONG
data and theoretical
calculated a_1, a_3

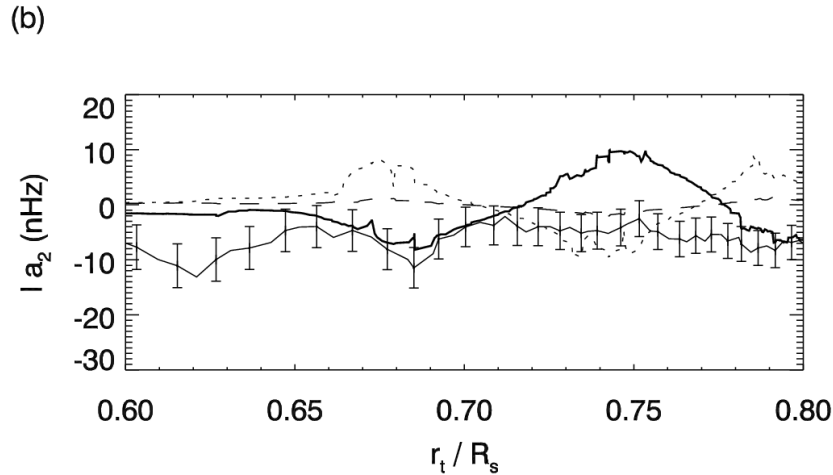
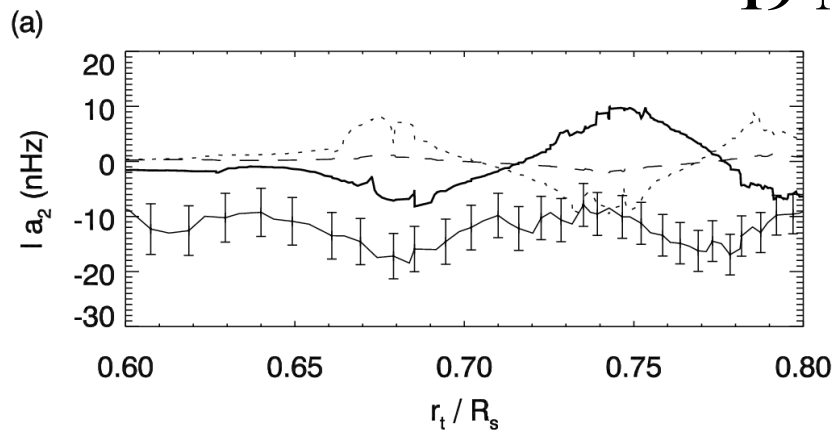
$$\Omega r \sin \theta = -w_1^0 \partial_\theta Y_1^0 - w_3^0 \partial_\theta Y_3^0$$

Observational splitting coefficients

- It is only the “nearly degenerate” modes we may hope to detect in observations.
- Use 7 different GONG data sets covering late descending part of cycle 22 – end of cycle 23
- Most contribution to a_2 comes from surface effects. But we are interested in modes with $r_t \sim 0.7$
- Multiplets sorted in order of increasing r_t and smoothed using a 100 point error weighted average.

Observational splitting coefficients

19-Nov-2002



18-Sep-2007

- Only way to put an upper limit on flow is to look for nearly degenerate modes in the data .
- **Problem:** running mean combines results with positive and negative a_q
- Separate into two groups on basis of theoretical results: $la_2 < -10$ nHz & $la_2 > 10$ nHz.

Observational splitting coefficients

- Consider modes with $la_2 < -10$ nHz and search for them in each GONG data set.
- Average la_2 from theory over available multiplets and error σ_2 in a_2 from observations.
- For Y_8^4 flow with $u_0 = 100$ ms⁻¹, $|la_2| \sim 22$ nHz (theory), $\sigma_2 \sim 14$ nHz (obs). Ruled out with a CL = $la_2 / \sigma_2 = 1.5$
- For Y_8^8 flow with $u_0 = 50$ ms⁻¹, $|la_2| \sim 28$ nHz (theory), $\sigma_2 \sim 13$ nHz (obs). Ruled out with a CL = $la_2 / \sigma_2 = 2.0$



Summary

- Meridional flow (Y_2^0) : theoretical coefficients la_2 are 7 times smaller than corresponding errors in observations. Not sensitive to no. of cells in radial direction. Impossible to detect return flow by this method.
- Meridional flow (Y_8^0) and giant cells (Y_8^8, Y_8^4): “Nearly degenerate” modes with $r_t \sim 0.7$ have large values of la_2 . Can hope to detect in observations.
- However we do not find any clear signal around $r_t \sim 0.7$, so we put upper limits on strength of flows by comparing la_2 (theory) with corresponding errors (observations).